

Braneworld Black Rings

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Plan of Talk

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Introduction

- Brane world scenario requires generic gravitational configurations on the brane to arise from a four dimensional bulk. The investigation of black hole configurations in this context has been an exciting aspect of study
- Chamblin *et al* ^{*} described a Schwarzschild black hole as a 3-brane intercept of a bulk black string in a Randall-sundrum brane world. This was soon generalized to the case of a rotating black hole on the 3-brane[†]
- The bulk black string solution was generalized to n dimensional RS braneworld with a single transverse AdS direction[‡] where it intercepted the (n-1)brane in a Myers Perry black hole
- This work attempts to describe five dimensional black ring solutions in an RS braneworld scenario

^{*}A. Chamblin, S. Hawking, H. Reall, Phys. Rev. D 61, 065007 (2000)

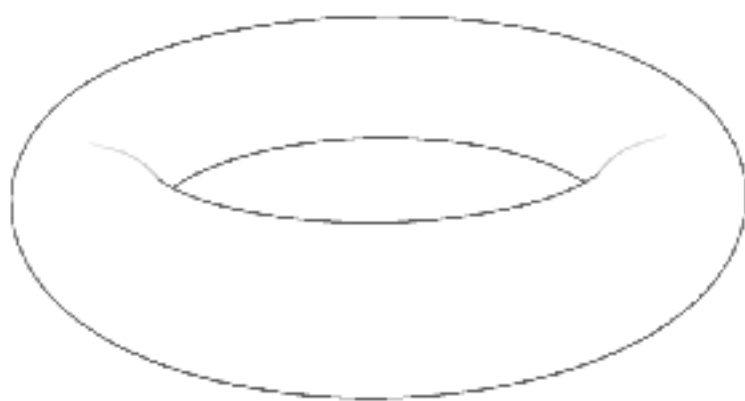
[†]M.S. Modgil, S. Panda and G.Sengupta, Mod. Phys. Lett A, 17 (2002) 1479

[‡]G.Sengupta, Int.J.Mod.Phys.D15 (2006) 171

Black Rings

- neutral rotating black rings* are stationary asymptotically flat solutions of five dimensional vacuum Einstein equations having a regular horizon of topology $S^1 \times S^2$.
- for a certain range of mass and angular momentum two black ring solutions can coexist with a rotating black hole solution. Thus, uniqueness doesn't hold in five dimensions.

*R. Emparan and H.S.Reall, Phys. Rev. Lett. 88 (2002)



Black ring cartoon. The azimuthal angle of S^2 is suppressed.

Co-ordinate system review

- $SO(4)$ has Cartan subgroup $U(1)^2$ thus implying two independent rotations.
- the four spatial co-ordinates can be paired as

$$\begin{aligned} x^1 &= r_1 \cos \phi & x^3 &= r_2 \cos \psi \\ x^2 &= r_1 \sin \phi & x^4 &= r_2 \sin \psi. \end{aligned} \quad (1)$$

with the corresponding flat space metric

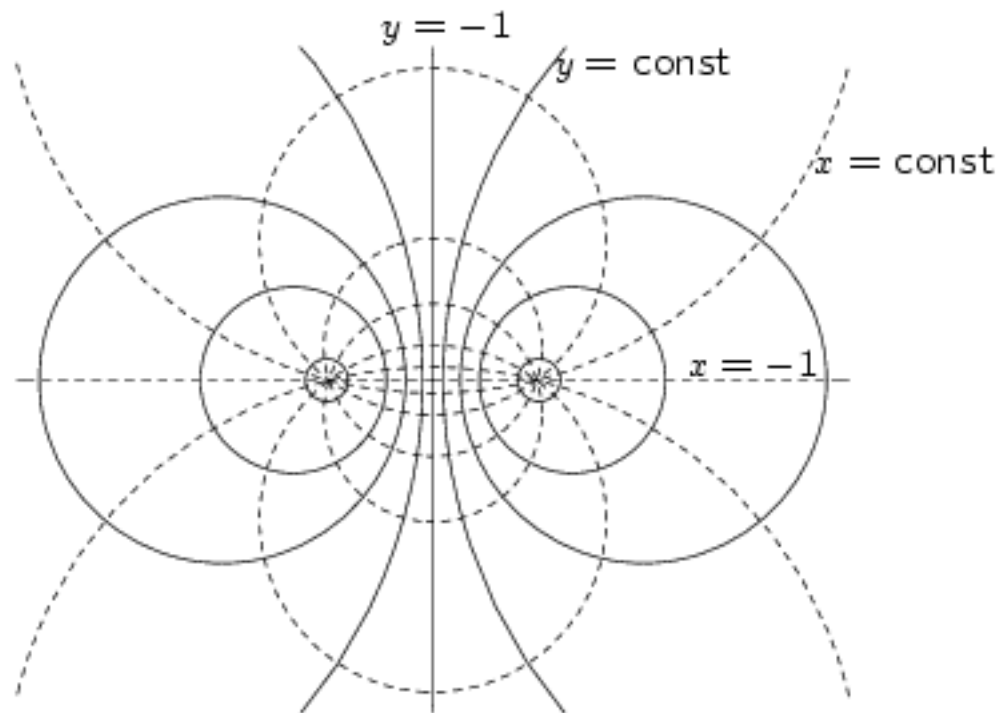
$$d\mathbf{x}_4^2 = dr_1^2 + r_1^2 d\phi^2 + dr_2^2 + r_2^2 d\psi^2. \quad (2)$$

- *adapted co-ordinates* (x, y, ψ, ϕ) foliate flat space by the equipotential surfaces of the field created by a source resembling the black ring extended along (x^3, x^4) plane and rotating along ψ direction.

$$r_1 = R \frac{\sqrt{1-x^2}}{x-y}, \quad r_2 = R \frac{\sqrt{y^2-1}}{x-y}. \quad (3)$$

$$-\infty \leq y \leq -1, \quad -1 \leq x \leq 1$$

- asymptotic infinity is the limit $x, y \rightarrow -1$
- $y = -\infty$ or $r_1 = 0, r_2 = R$ is the location of the ring source
- $y = -1$, the fixed 'point' of ψ rotation, is the whole of (r_1, ϕ) plane.
- $x = 1$ and $x = -1$, the fixed 'points' of ϕ rotation are, respectively, $r_2 \leq R$ and $r_2 > R$ regions of (r_2, ψ) plane.



Ring coordinates for flat four-dimensional space, on a section at constant ϕ and ψ (and $\phi + \pi$, $\psi + \pi$). Dashed circles correspond to spheres at constant $|x| \in [0, 1]$, solid circles to spheres at constant $y \in [-\infty, -1]$. Spheres at constant y collapse to zero size at the location of the ring of radius R , $y = -\infty^*$.

flat space metric in **ring co-ordinates**

$$d\mathbf{x}_4^2 = \frac{R^2}{(x-y)^2} \left[(y^2 - 1)d\psi^2 + \frac{dy^2}{y^2 - 1} + \frac{dx^2}{1 - x^2} + (1 - x^2)d\phi^2 \right]. \quad (4)$$

- constant y sections have ring like topology $S^1 \times S^2$
- the S^2 's are metrically round spheres for flat space.

*R. Emparan, H.S Reall Class.Quant.Grav. 23 (2006) R169

Black ring Metric

The metric of a neutral rotating black ring is

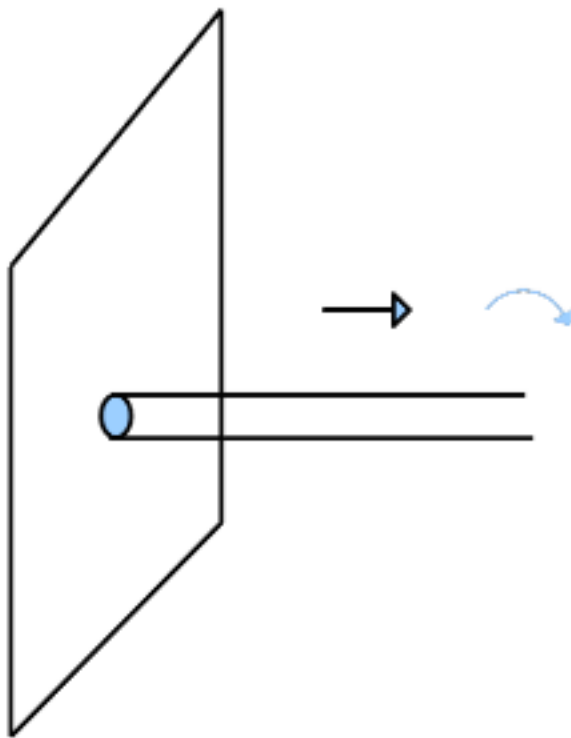
$$ds^2 = -\frac{F(y)}{F(x)} \left(dt - C R \frac{1+y}{F(y)} d\psi \right)^2 + \frac{R^2}{(x-y)^2} F(x) \left[-\frac{G(y)}{F(y)} d\psi^2 - \frac{dy^2}{G(y)} + \frac{dx^2}{G(x)} + \frac{G(x)}{F(x)} d\phi^2 \right] \quad (5)$$

$$F(\xi) = 1 + \lambda\xi, \quad G(\xi) = (1 - \xi^2)(1 + \nu\xi) \quad C = \sqrt{\lambda(\lambda - \nu) \left(\frac{1+\lambda}{1-\lambda} \right)}.$$

$$0 < \nu \leq \lambda < 1 \quad (6)$$

- three Killing isometries $\partial_t, \partial_\phi, \partial_\psi$
- ν signifies shape of the rings
- λ/ν measures speed of rotation along ψ , i.e S^1
- $y = -1/\lambda$ section is the ergosurface
- $y = -1/\nu$ section is the black ring horizon
- $y = -\infty$ is a spacelike curvature singularity inside the horizon

- black ring tension and gravitational self attraction are **internally balanced** by the centrifugal force. This implies there are only two independent parameters of black ring, R and ν . Also, it means angular momentum is bounded from below.
- black ring of large radius R is like a Schwarzschild **black string boosted** along the translation invariant direction and periodically identified with a period $2\pi R$. The pressure of the boosted string is zero



Black ring geodesics

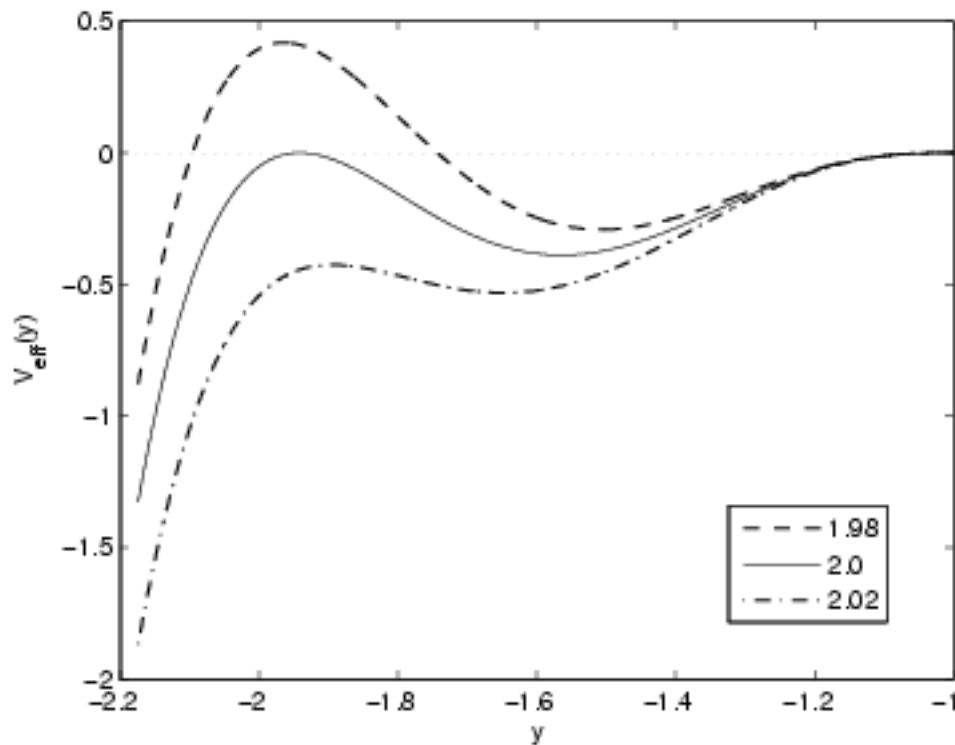
- **equatorial plane:** we consider geodesics restricted to the plane of rotation of the black ring outside the horizon, *i.e.* $r_2 > R$ or, equivalently, the $x = -1$ section in the ring co-ordinates.

$$\dot{y}^2 + V_{eff}(y; E, \Psi) = 0 \quad (7)$$

- asymptotic behaviour ($y \rightarrow -1$)

$$V_{eff}(y; E, \Psi) \rightarrow -\frac{2(1-\nu)}{R^2(1+\lambda)}|1+y|^3(E^2 - m^2), \quad (8)$$

- unbounded timelike geodesics only when $E^2 > m^2$



Effective potential plot. Three values of E are as indicated in the box. The constants $m = R = 1^*$.

It is apparent that there are **no stable bound orbits**, as expected for central potentials in more than four dimensions.

*Anurag Sahay, Gautam Sengupta JHEP06(2007)006

Braneworld Black Ring

- AdS6 in conformally flat co-ordinates

$$ds^2 = \frac{l^2}{z^2} [\eta_{\mu\nu} dx^\mu dx^\nu + dz^2] \quad (9)$$

- 5 + 1 dimensional [RS2 braneworld](#) with a single transverse AdS direction and a reflection symmetric 4-brane obtained from AdS6 by slicing, orbifolding and then pushing the regulator brane to AdS horizon.
- rotating black string in ADS6 as the bulk solution intercepting the 4-brane in a five dimensional black ring.

$$ds^2 = \frac{l^2}{(z_0 + |w|)^2} \left[dw^2 - \frac{F(y)}{F(x)} \left(dt - CR \frac{1+y}{F(y)} d\psi \right)^2 \right. \\ \left. + \frac{R^2}{(x-y)^2} F(x) \left(-\frac{G(y)}{F(y)} d\psi^2 - \frac{dy^2}{G(y)} + \frac{dx^2}{G(x)} + \frac{G(x)}{F(x)} d\phi^2 \right) \right] \quad (10)$$

- ADM mass and angular momentum on the brane are scaled by the warp factor

$$M_* = \left(\frac{l}{z_0}\right)^2 M, \quad J_* = \left(\frac{l}{z_0}\right)^3 J. \quad (11)$$

- curvature squared diverges at the black ring singularity on the brane and at the AdS horizon ($z \rightarrow -\infty$) as well

$$\begin{aligned} R_{jklm}R^{jklm} &= \frac{1}{l^4} \left[60 + \frac{6(1+\nu^2)^2\nu^2 Q(x,y)}{R^4(1+\nu^2+2\nu x)^6} z^4 (x-y)^4 \right] \\ &\sim \frac{z^4}{r^8} + \text{const.} \end{aligned} \quad (12)$$

- horizon singularity artifact of linearization(?)
check with the geodesics

Black 2-brane

- in the large R limit the rotating black string reduces to a bulk black 2-brane in AdS_6 boosted along the extended direction on the 4-brane and identified periodically.

$$ds^2 = \frac{l^2}{z^2} \left[dz^2 + du^2 - \left(1 - \frac{r_0}{r}\right) dt^2 + \left(1 - \frac{r_0}{r}\right)^{-1} dr^2 + r^2 d\Omega_2^2 \right] \quad (13)$$

Braneworld Geodesics

- bulk geodesics that reach the AdS horizon inform the nature of singularity there
- we consider geodesics which on the 4-brane are restricted to the equatorial plane of the black ring.

$$\left(\frac{dy}{d\rho}\right)^2 + \frac{z^4}{l^4} g^{yy} \left(\frac{l^2 m^2}{z_1^2} + g^{tt} E^2 - g^{t\psi} E \Psi + g^{\psi\psi} \Psi^2 \right) = 0 \quad (14)$$

$$z \rightarrow -\infty \text{ as } \rho \rightarrow 0^-$$

- both timelike and null geodesics when projected onto the 4-brane by scaling out z dependence behave as timelike geodesics on the brane with proper time γ

$$\left(\frac{d\tilde{y}}{d\gamma}\right)^2 + V_{eff}(\tilde{y}; \tilde{E}, \tilde{\Psi}) = 0 \quad (15)$$

$$\gamma \rightarrow \infty \text{ as } \rho \rightarrow 0^-$$

- late time behaviour of projected geodesics on the brane equivalent to near horizon behaviour of bulk geodesics
- no stable bound geodesic on the brane so only the bulk geodesics with unbounded projections reach the horizon
- along such geodesics the curvature squared remains finite thus indicating the presence of p-p curvature singularity
- explicit determination necessitates obtaining curvature components in a pp orthonormal frame.. intractable so far!

Conclusion

- we have described five dimensional black ring in a Randall-Sundrum braneworld perspective as the intercept on the 4-brane of a rotating black string solution in AdS_6
- this is motivated by the fact that for consistency, gravitational configurations on the brane, in particular black holes, arise from some higher dimensional bulk solutions.
- black rings and their generalizations like black di-rings, black saturn, doubly spinning black rings, etc, make for an interesting configuration to study from a bulk brane world perspective.