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Gravitational potential and tidal tensor.

The gravitation acceleration at any point can be expressed as the gradient of a potential $\phi(\vec{x}, t)$

$$\vec{a} = -\vec{\nabla} \phi(\vec{x}, t).$$

$$\text{and } \vec{F} = -m \vec{\nabla} \phi(\vec{x}, t)$$

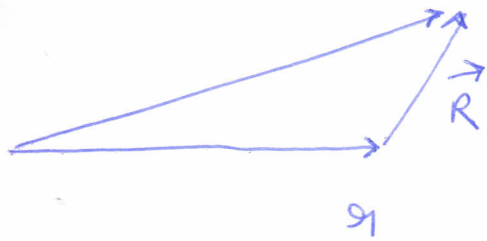
where m is the mass of the test particle located at \vec{x} .

$$\phi(\vec{x}, t) = -G \int \frac{\rho(\vec{x}', t) d^3x'}{|\vec{x} - \vec{x}'|}$$

$$\nabla^2 \phi(\vec{x}, t) = 4\pi G \rho(\vec{x}, t).$$

Poisson equation.

Tidal force.



$$R \ll r.$$

$$\vec{F}_T(\vec{r}, \vec{R}, t) = \vec{F}(\vec{r} + \vec{R}, t) - \vec{F}(\vec{r}, t)$$

$$\odot F_{T,i} = -m \left[\phi_{,i}(\vec{r} + \vec{R}, t) - \phi_{,i}(\vec{r}, t) \right]$$

$$\approx -m \phi_{,ij} R_j \quad \text{where the derivative}$$

is at r .

Consider the situation where there is a point mass at \vec{r} . This causes an isotropic inwards force which is not tidal. We have to subtract this out.

$$F_T = -m \left[\phi_{,ij} - \frac{1}{3} \delta_{ij} \nabla^2 \phi \right] R_j$$

we define the tidal tensor

$$T_{ij} = - \left[\phi_{,ij} - \frac{1}{3} \delta_{ij} \nabla^2 \phi \right]$$

example along z axis from a mass

M.



$$\phi(r) = - \frac{GM}{r}$$

$$\phi_{,i} = \frac{GM}{r^3} r_i$$

$$\phi_{,ij} = \frac{GM r^2 \delta_{ij} - 3 r_i r_j}{r^5}$$

$$T_{ij} = - \frac{GM}{r^5} [r^2 \delta_{ij} - 3 r_i r_j]$$

$$T_{11} = - \frac{GM}{r^3} = T_{22}$$

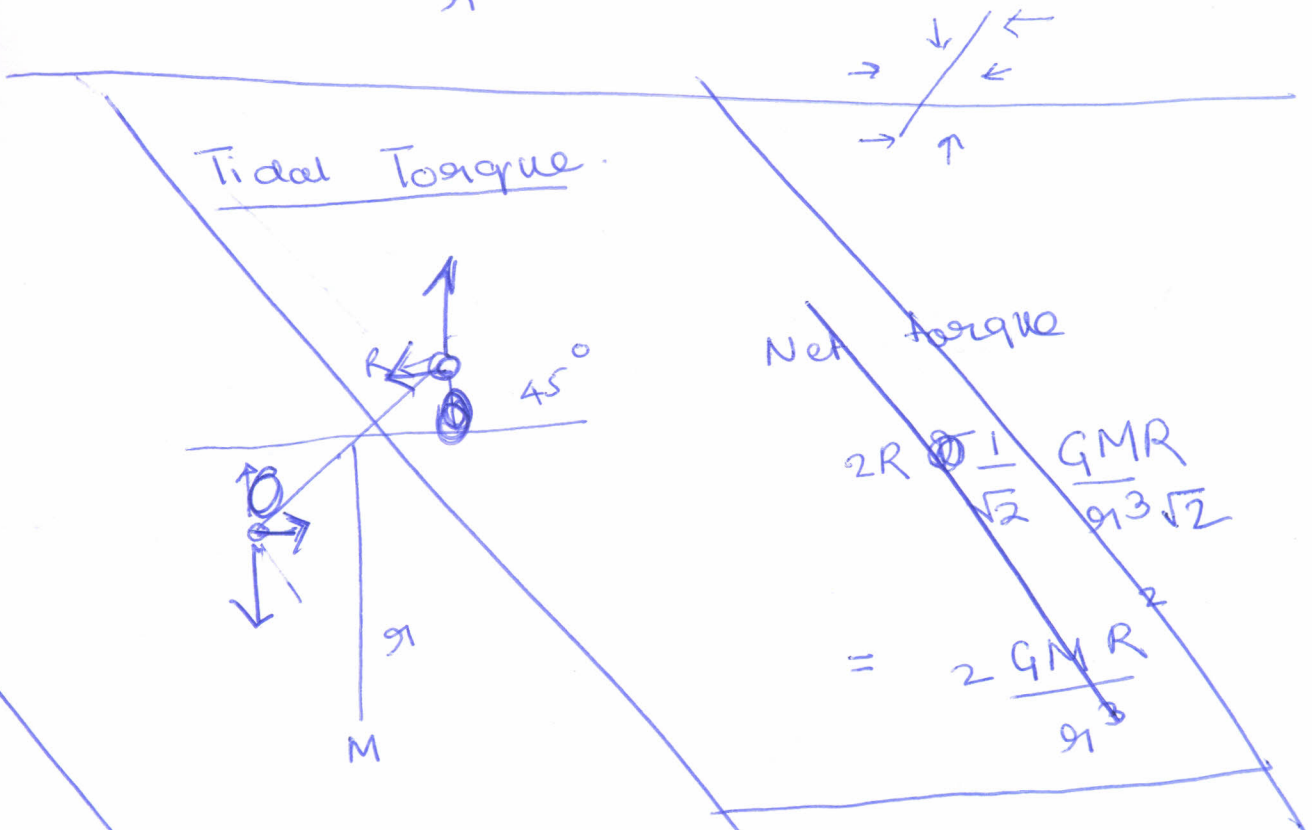
$$T_{33} = \frac{2GM}{r^3}$$

$$\begin{pmatrix} F_{Tx} \\ F_{Ty} \\ F_{Tz} \end{pmatrix} = \frac{GM}{r^3} \begin{pmatrix} -1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & 2 \end{pmatrix} \begin{pmatrix} R_x \\ R_y \\ R_z \end{pmatrix}$$

$$F_{Tz} = \frac{2GM}{r^3} R_z ; F_{Tx} = -\frac{GM}{r^3} R_x$$

$$F_{Ty} = -\frac{GM}{r^3} R_y$$

as discussed earlier -



$$\begin{aligned} \text{Torque} &= \frac{R}{\sqrt{2}} \times \frac{GM R}{r^3 \sqrt{2}} + \frac{2GM}{r^3} \frac{R}{\sqrt{2}} \frac{R}{\sqrt{2}} \\ &= \frac{GM R^2}{r^3} \times \frac{3}{2} \end{aligned}$$

in general.

$$\tau_i = \epsilon_{ijk} \int \rho(\mathbf{r}) (\mathbf{r} - \mathbf{r}_c)_j (\mathbf{r} - \mathbf{r}_c)_k d^3r$$

around center of mass.

$$\tau_i = \epsilon_{ijk} \int \rho(\mathbf{r} - \mathbf{r}_c)_j \cancel{\rho(\mathbf{r})} F_k(\mathbf{r}) d^3r$$

$$= \epsilon_{ijk} \int (\mathbf{r} - \mathbf{r}_c)_j \left[-\rho(\mathbf{r}) \left\{ \phi_{,k}(\mathbf{r}_c) + \phi_{,ke}(\mathbf{r}_c) (\mathbf{r} - \mathbf{r}_c)_e \right\} \right] d^3r$$

$$= \epsilon_{ijk} \left[\int (\mathbf{r} - \mathbf{r}_c)_j \rho(\mathbf{r}) d^3r \right] \left[-\phi_{,k}(\mathbf{r}_c) \right]$$

$$- \epsilon_{ijk} \phi_{,ke}(\mathbf{r}_c) \int d^3r (\mathbf{r} - \mathbf{r}_c)_j (\mathbf{r} - \mathbf{r}_c)_e \rho(\mathbf{r})$$

The first integral is zero, definition of

C.M.
Second integral.

$$= \epsilon_{ijk} \left[T_{ke} + \frac{1}{3} \delta_{ke} \nabla^2 \phi \right] I_{je}$$

we have $\epsilon_{ijk} I_{jk} = 0$

I_{jk} is the symmetric moment of inertia tensor.

① In a similar way.

$$\zeta_i = \epsilon_{ijk} T_{ke} \left[I_{je} - \frac{\delta_{jle}}{3} I_{mm} \right]$$

←—————→

traceless part of inertia tensor.

arises due to deviation from spherical symmetry quadrupole tensor. — this is the mass tensor.

$$Q_{ke} = \int \rho(r) \left[(r-r_c)_e (r-r_c)_k - \frac{\delta_{ke}}{3} (r-r_c)^2 \right] d^3r.$$

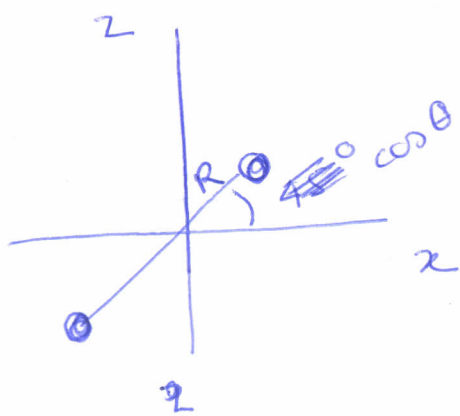
This is the tidal torque.

arises only if there is a mass quadrupole moment.

Question How does it align the quadrupole relative to the direction of the local gravitation acceleration.

Final

$$\zeta_i = \epsilon_{ijk} A_{je} T_{ke}$$



$$Q = R^2 \begin{pmatrix} \cos^2 \theta & 0 & \cos \theta \sin \theta \\ 0 & 0 & 0 \\ \cos \theta \sin \theta & 0 & \sin^2 \theta \end{pmatrix}$$

$$= \frac{2}{3} R^2 \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

$$Q_{jk} = 2 R^2 \left[n_j n_k - \frac{\delta_{jk}}{3} \right]$$

$$T_{mj} = - \frac{GM}{r^3} \left[3 \hat{z}_m \hat{z}_j - \delta_{mj} \right]$$

$$\epsilon_{imk} T_{mj} Q_{jk} = - \frac{GM R^2}{r^3} \left[\hat{z}_m \hat{n}_j + \frac{\hat{z}_m \hat{n}_k}{3} + \left(\hat{z}_m \hat{n} \right) \right]$$

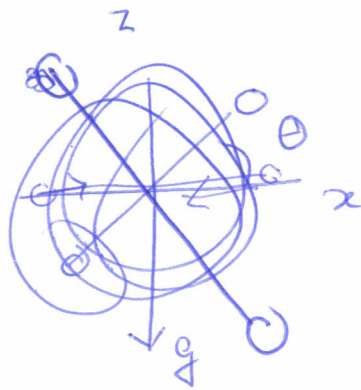
$$= - \frac{GM R^2}{r^3} \epsilon_{imk} \hat{z}_m \hat{n}_k \left(\hat{z} \cdot \hat{n} \right)$$

$$= - \frac{2GM R^2}{r^3} \epsilon_{ijk} \hat{n}_k (\hat{n}_i \hat{n}_j)$$

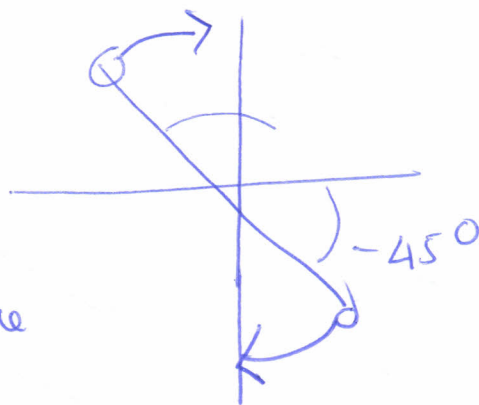
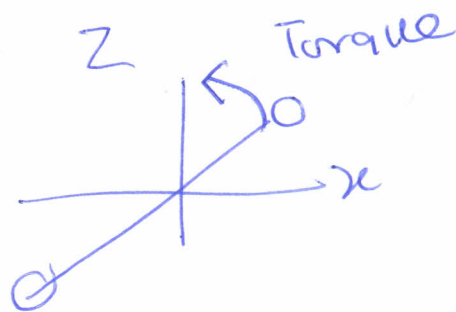
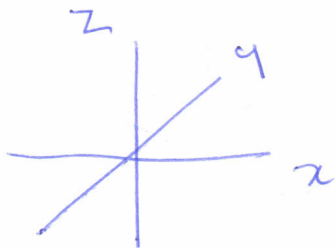
Let us take \hat{n} in $x-z$ plane

$$\hat{n}_1 = \cos \theta$$

$$\tau_2 = - \frac{2GM R^2}{r^3} \epsilon_{231} \cos \theta \sin \theta$$



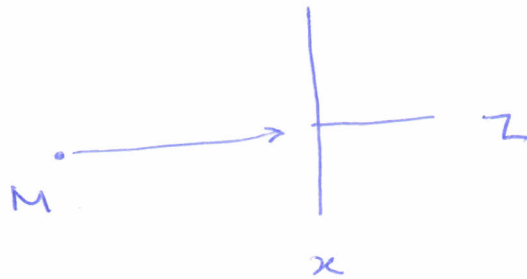
Torque ~~is~~ outwards



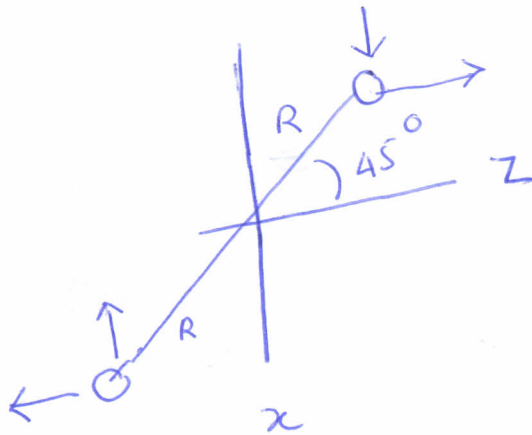
align along the direction of $\nabla \phi$.

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Torque on a dumbbell from a mass a distance r away



y axis points inwards



$$(\text{Torque})_y = \frac{2GMm}{r^3} 2 \times \frac{R}{\sqrt{2}} \times \frac{R}{\sqrt{2}}$$

$$+ \frac{GM}{r^3} 2 \times \frac{R}{\sqrt{2}} \times \frac{R}{\sqrt{2}}$$

$$= 3GMm \frac{R^2}{r^3}$$

The torque aligns the dumbbell with \vec{g} .