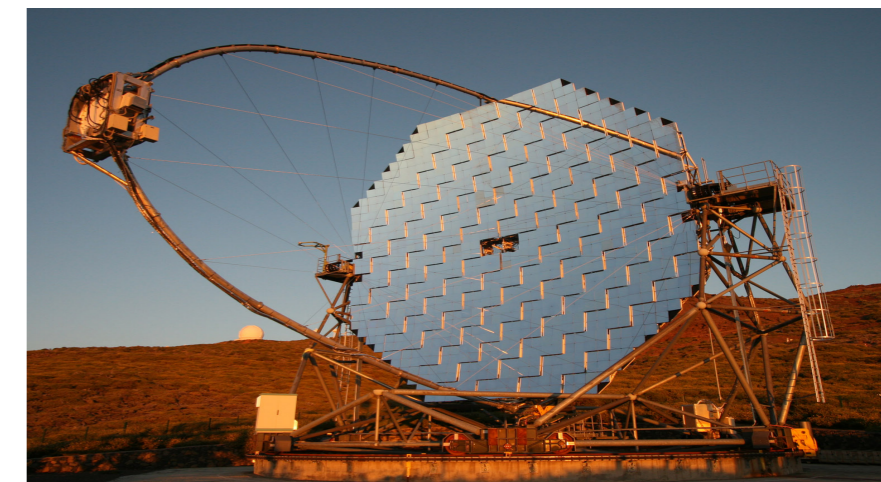
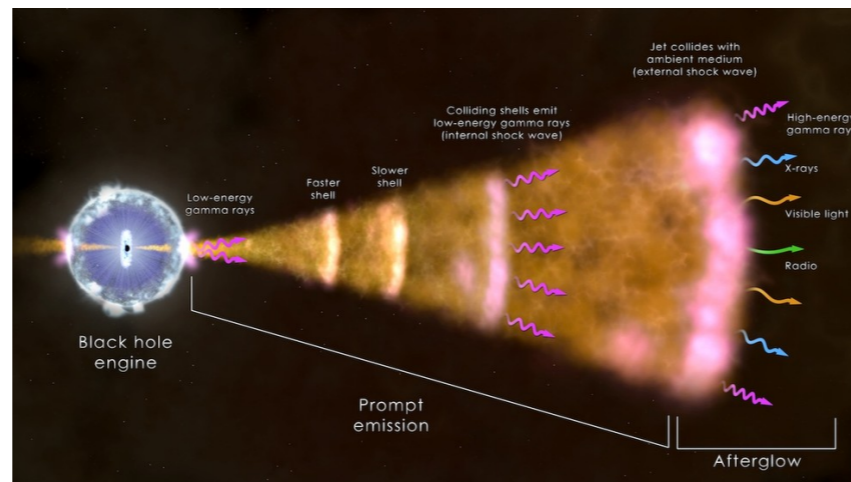
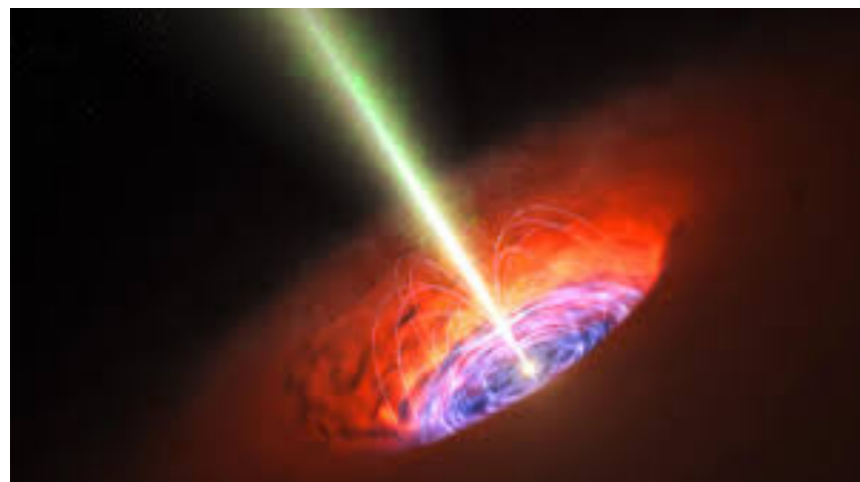
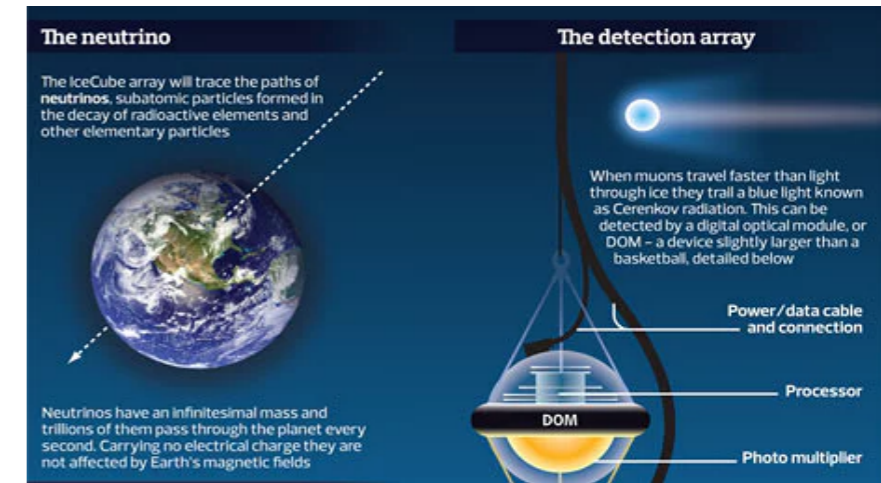
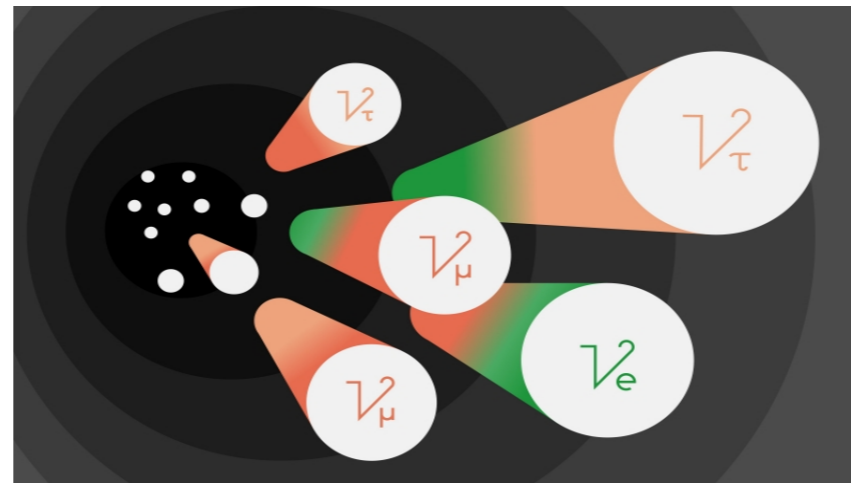
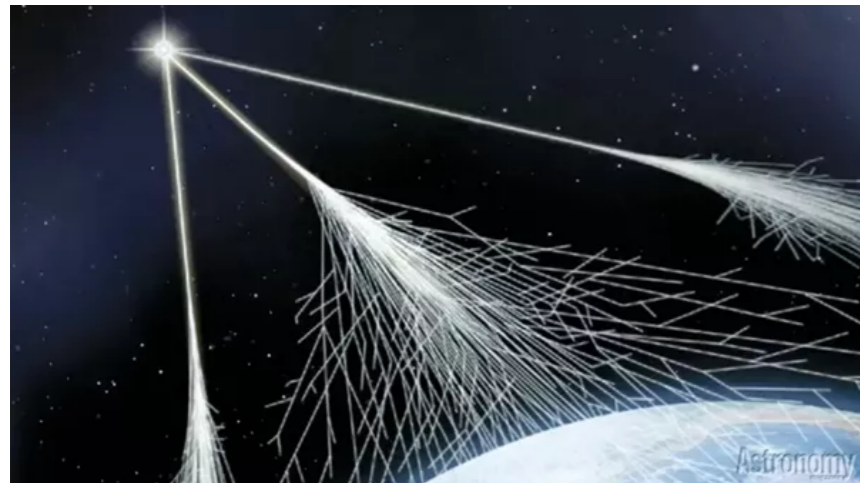


# High Energy Universe

## Cosmic Rays & Neutrinos & Gamma Rays



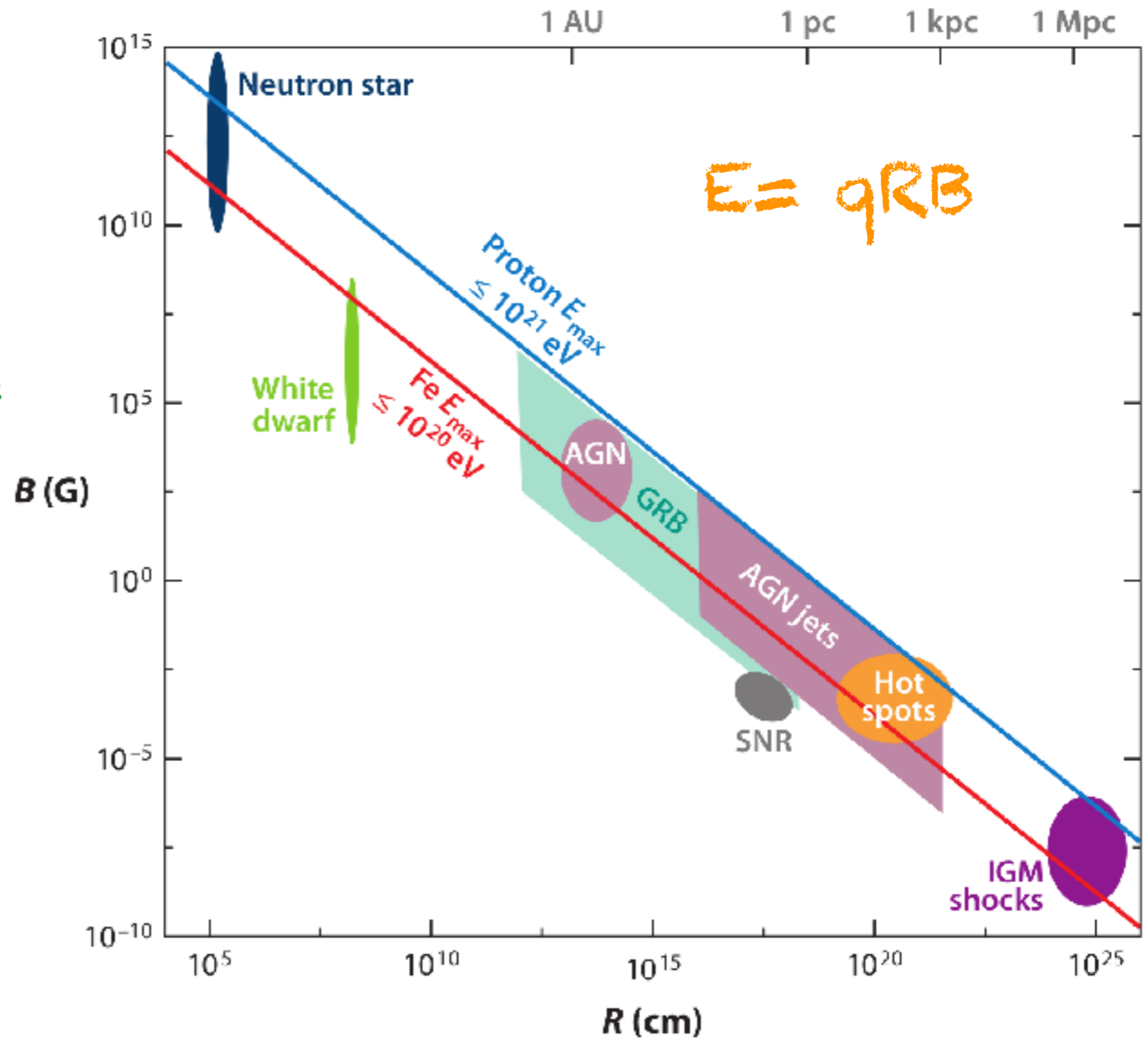
Debanjan Bose [debaice@gmail.com]

Lecture 2 [11.04.2019]

# Possible Sources of Cosmic Rays

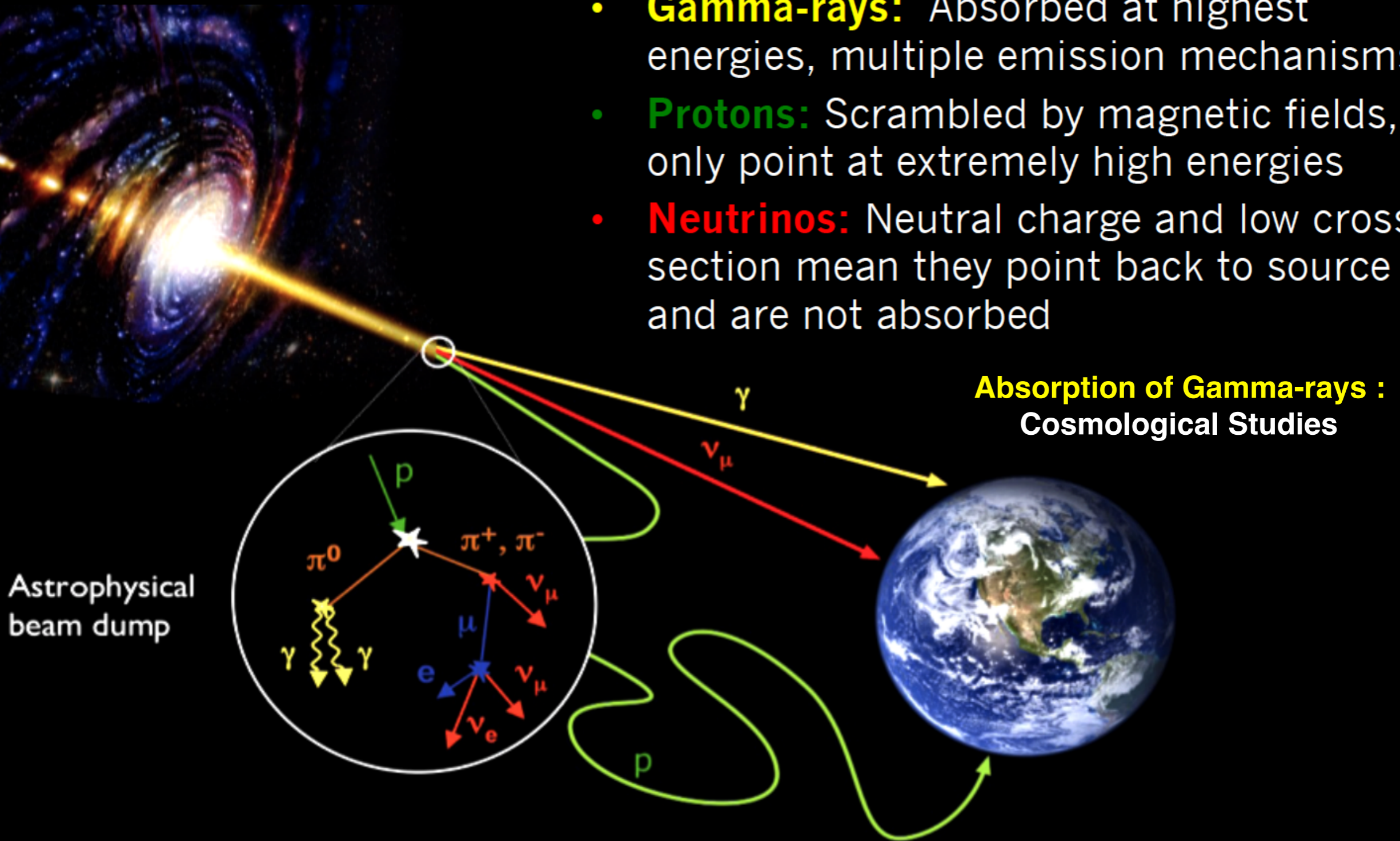
# General Principle of Acceleration & Probable Sources (Hillas Plot)

Magnetic fields themselves do not work therefore can not be directly responsible for accl. But changing magnetic field leads to inductive electric field

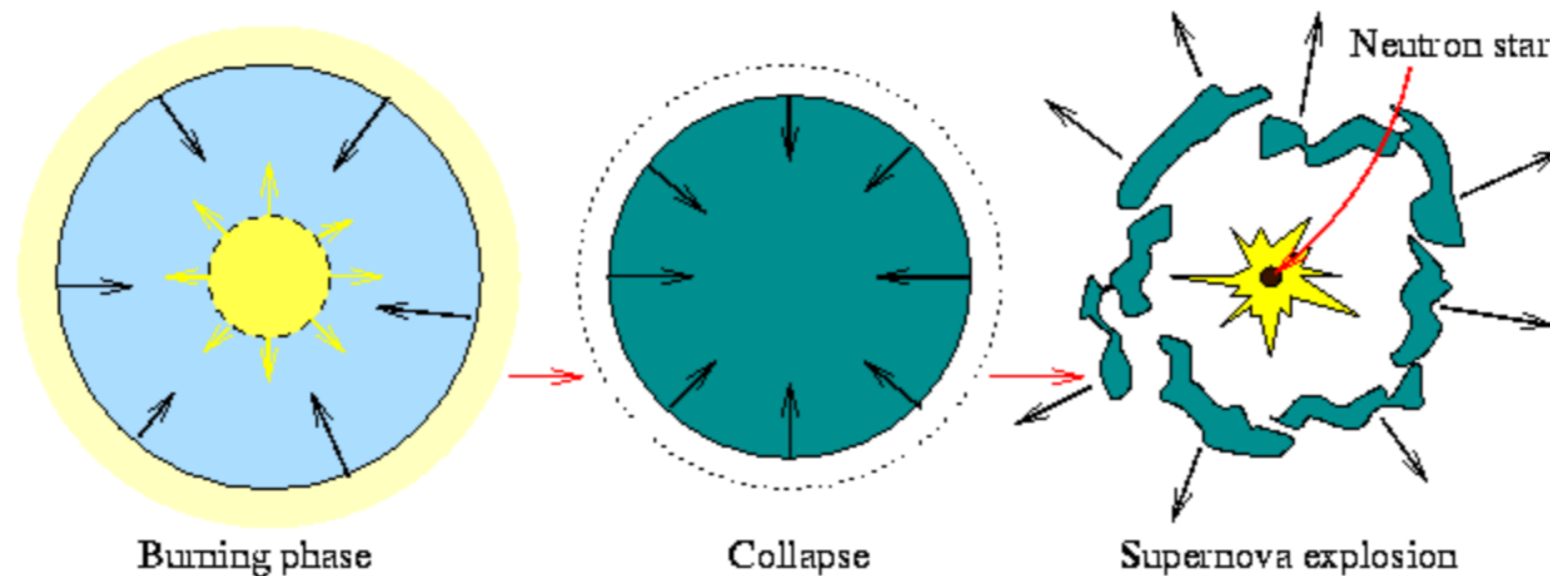


# Cosmic Messengers @ High Energies

- **Gamma-rays:** Absorbed at highest energies, multiple emission mechanisms
- **Protons:** Scrambled by magnetic fields, only point at extremely high energies
- **Neutrinos:** Neutral charge and low cross-section mean they point back to source and are not absorbed



# Neutron Stars



*Density  $\sim 10^{14} \text{g/cm}^3$ , Radius  $\sim 10 \text{km}$ , Mass  $\sim 1.5M_{\odot}$*

Once the Star completes all its nuclear burning - core collapses under the gravitational force produced by its own mass : **Supernova Explosion**

SN explosion decreases star radius by 5 orders of magnitude, increases magnetic field strength to  $10^8$  to  $10^{12}$  G : conservation of angular momentum and magnetic flux

Shortest known period of a pulsar is 1 ms. In that time light can travel 300 km. So emission region must be smaller than that.. compact region.

# SN - Source of Galactic CRs (Upto Knee)

Energy Density of CRs in our Galaxy  $1 \text{ eV/cm}^3$

Milky way has radius 20 kpc and height 200 pc

Total CR energy in the galaxy  $E_{G,CR} = \pi R^2 h \rho = 5 \times 10^{54} \text{ ergs}$

On an avg a CR spends  $10^{14}$  secs in the Galaxy confined to its magnetic field  $3 \mu\text{G}$ . Therefore there should be astronomical objects which has accleration power  $10^{40} \text{ erg/sec}$ .

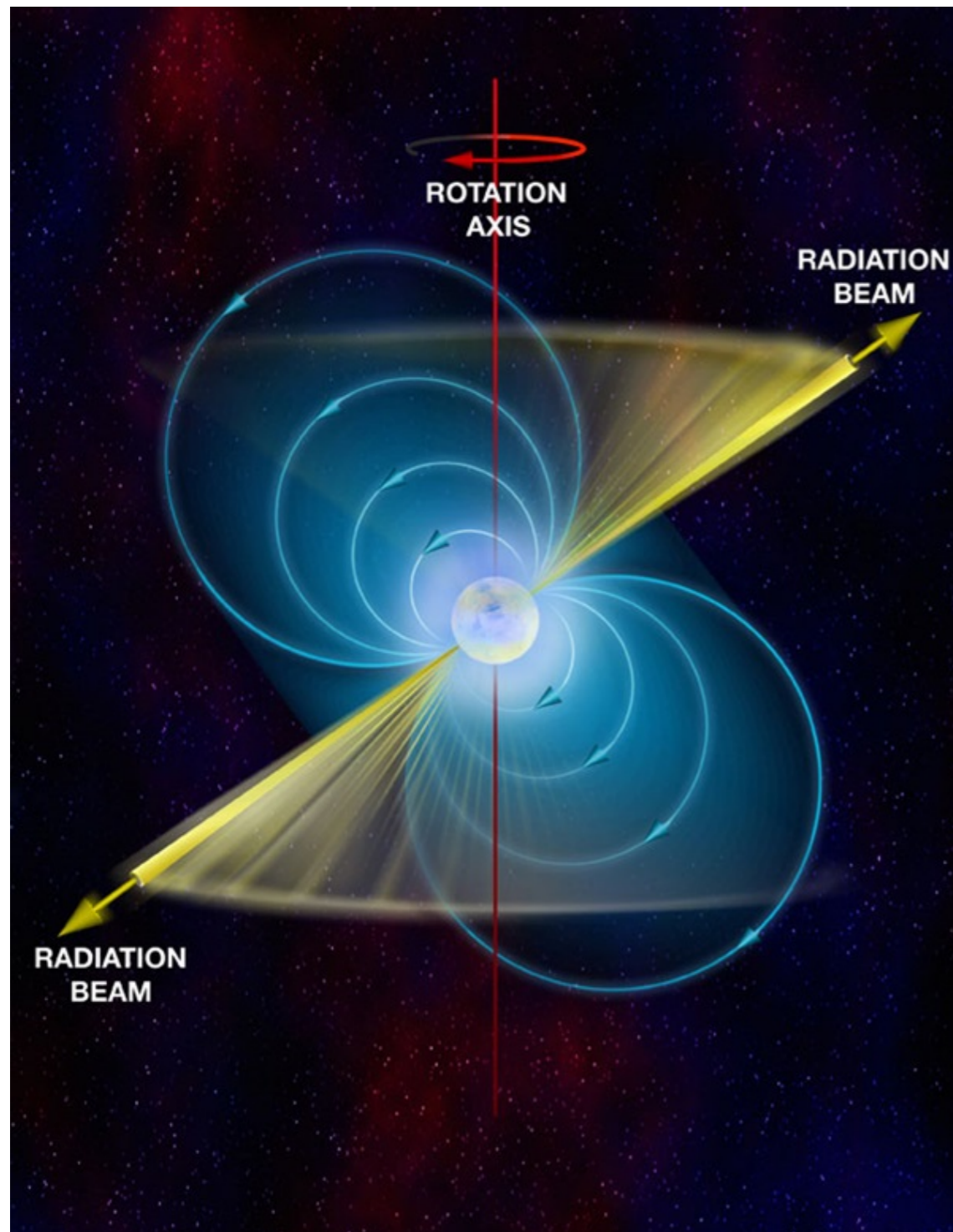
Avg total energy in SN explosion  $E_{SN} = 10^{51} \text{ erg}$

If 0.01% transferred to CR particles by Fermi Acceleration

And if there is 1 SN every 30 years then accleration power

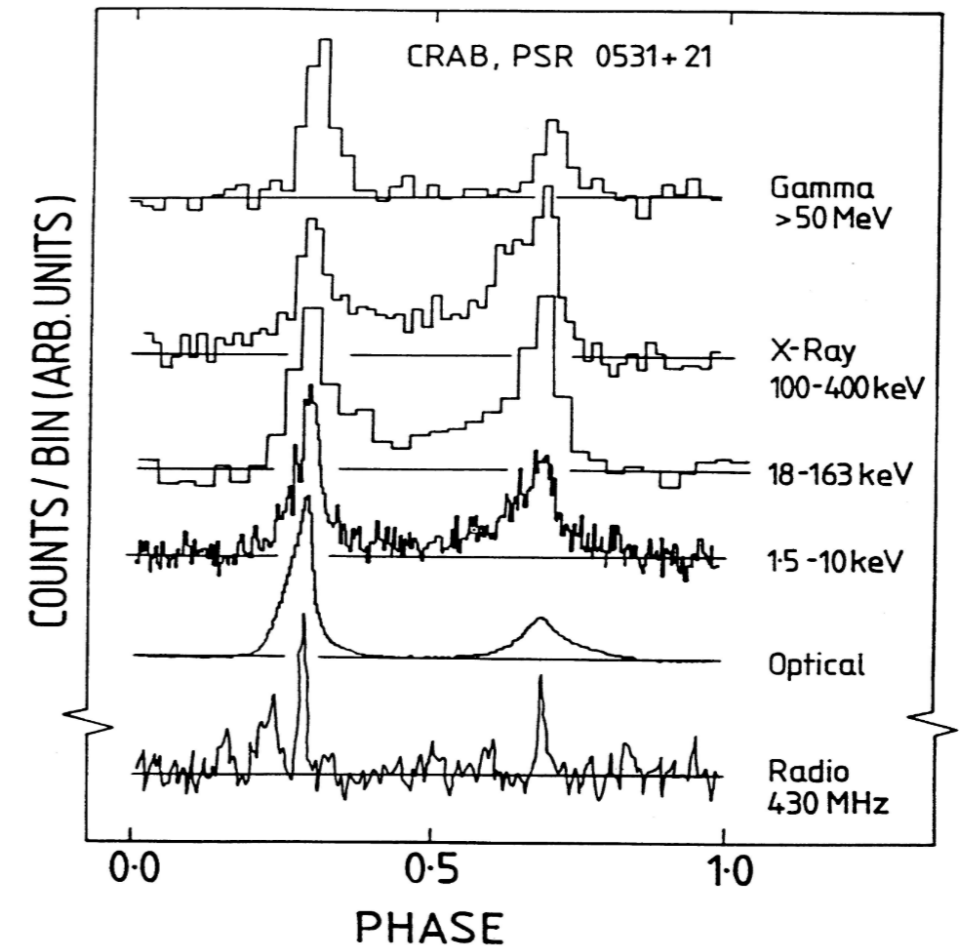
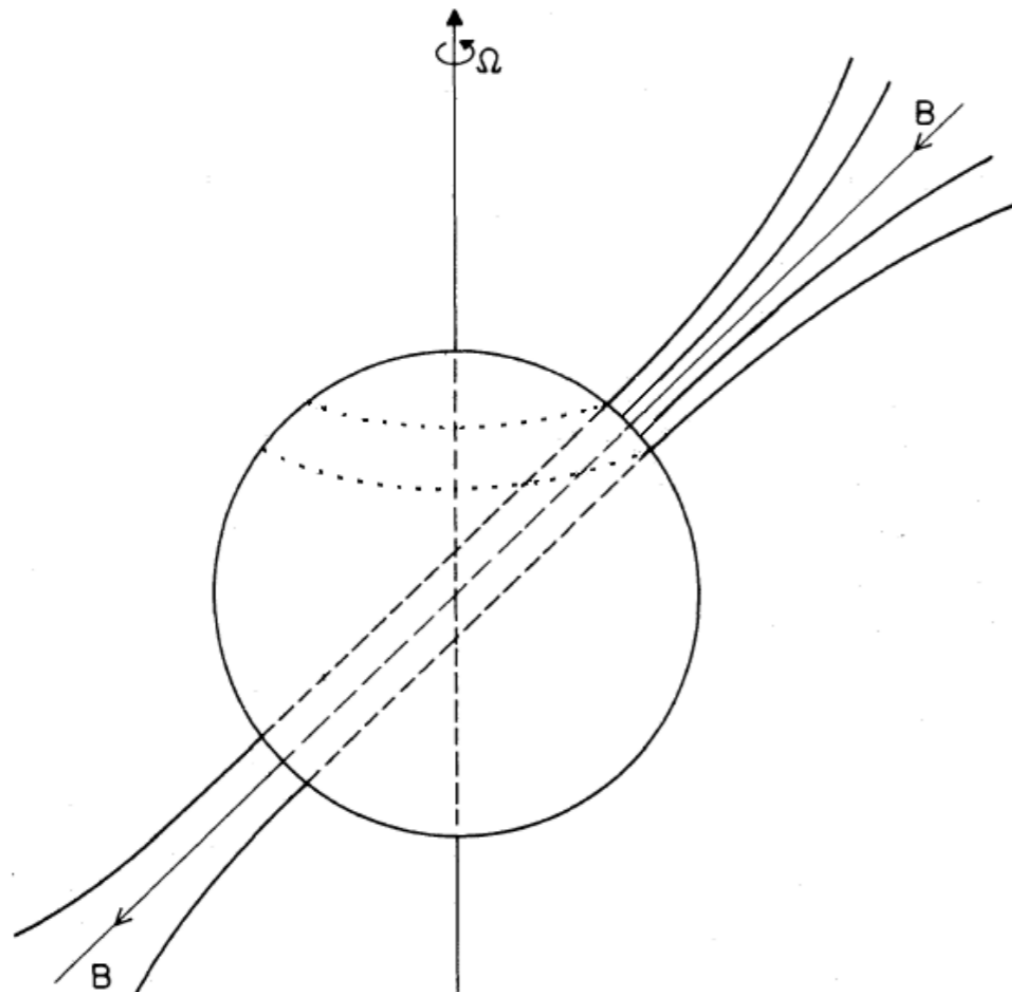
$$P = 0.01 E_{SN}/30 = 10^{40} \text{ erg/sec}$$

# Pulsars

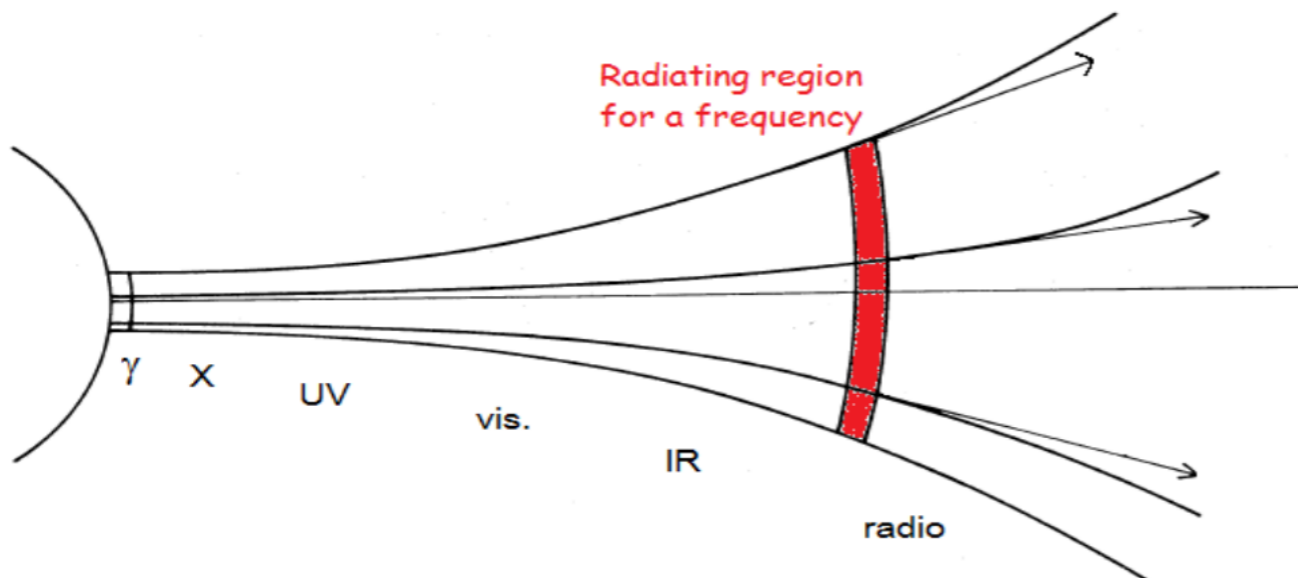


- A Pulsar - A pulsating Object [Pulsating EM radiation]
- Pulsar Period 1 ms to 10 secs
- Rapidly rotating Neutron Stars (NS)
- Most of the pulsars detected in radio emission : electrons & positrons are accelerated along the magnetized field lines of the NS

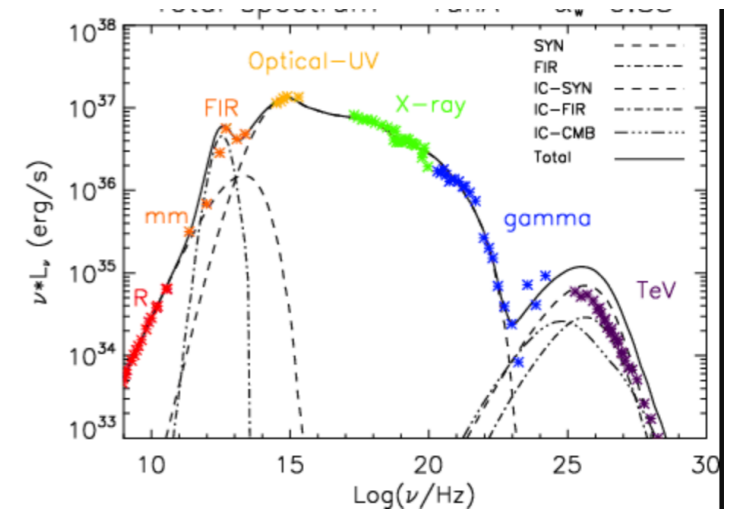
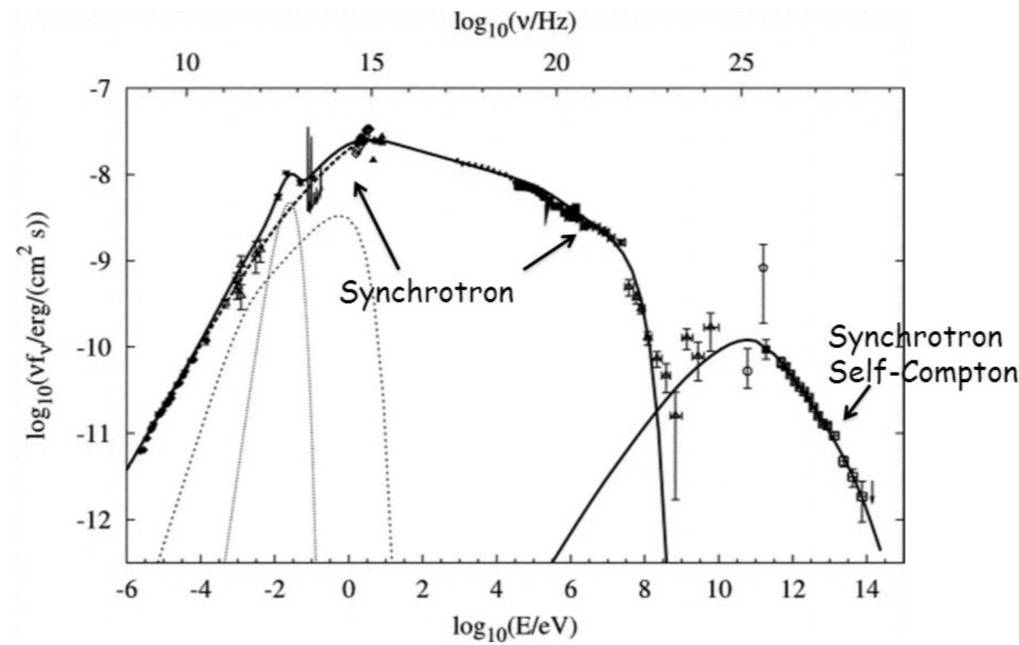
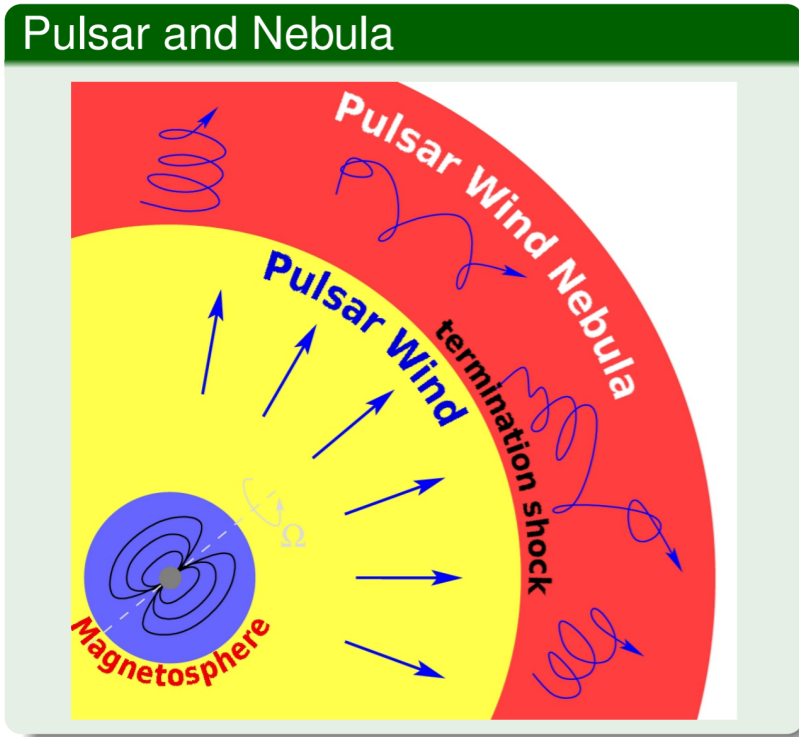
# HE Pulsed Emission from Crab Pulsar



Due to rapid rotation of very strong magnetic field huge electric field is induced  $10^{11}$  V/m. And they pull  $e^\pm$  away from NS surface overcoming gravitational pull. These particles move along magnetic lines and emit EM radiation.



# HE Continuous Emission from Crab Pulsar



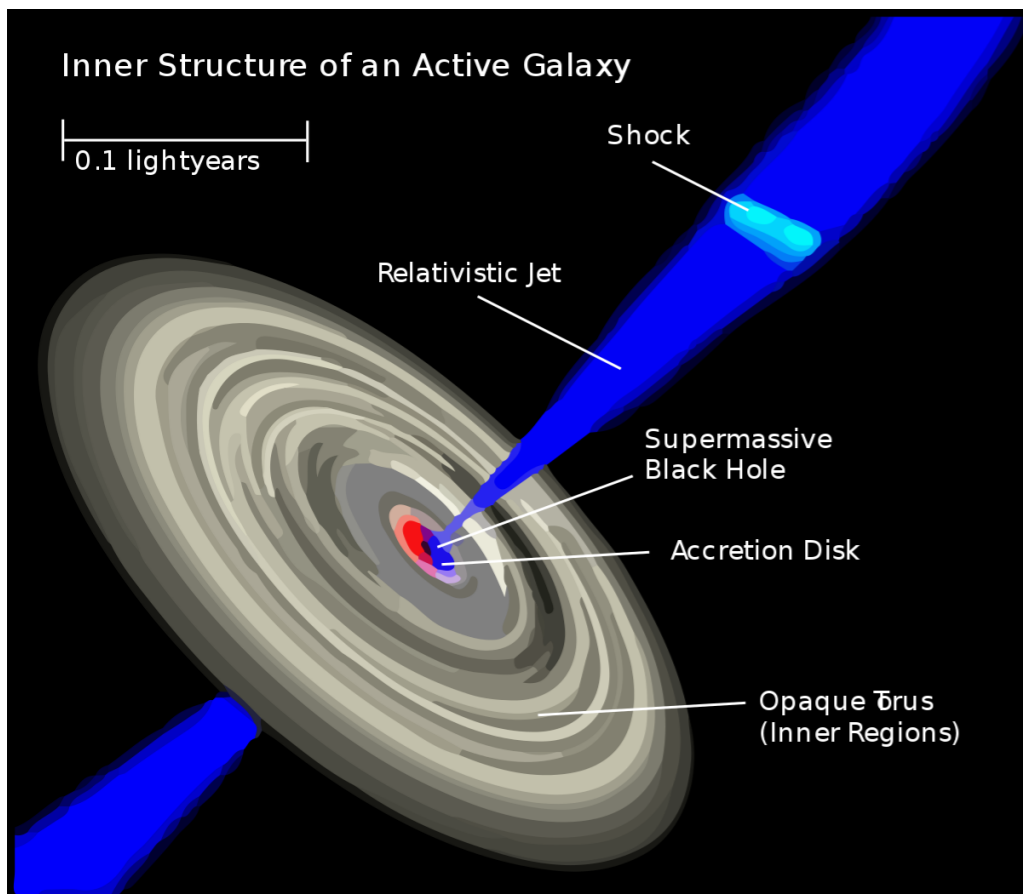
Ground based gamma-ray telescopes have detected both pulsed and continuous emission from Crab. Continuous emission comes from pulsar wind nebula. This continuous emission is used as candle to calibrate telescopes.

According to some models protons are also presents along with in the magnetized wind of relativistic plasma, p-p and p- $\gamma$  processes produce

$\gamma$  and  $\nu$

# Sources of Extra-Galactic Cosmic rays

## Active Galactic Nuclei (AGN)



Energy density of extra-galactic cosmic ray :

$$10^{36} \text{ erg Mpc}^{-3} \text{ s}^{-1}$$

In order to maintain this density total power requirement for a source with spectrum of  $E^{-2}$  would be :

$$2 \times 10^{37} \text{ erg Mpc}^{-3} \text{ s}^{-1}$$

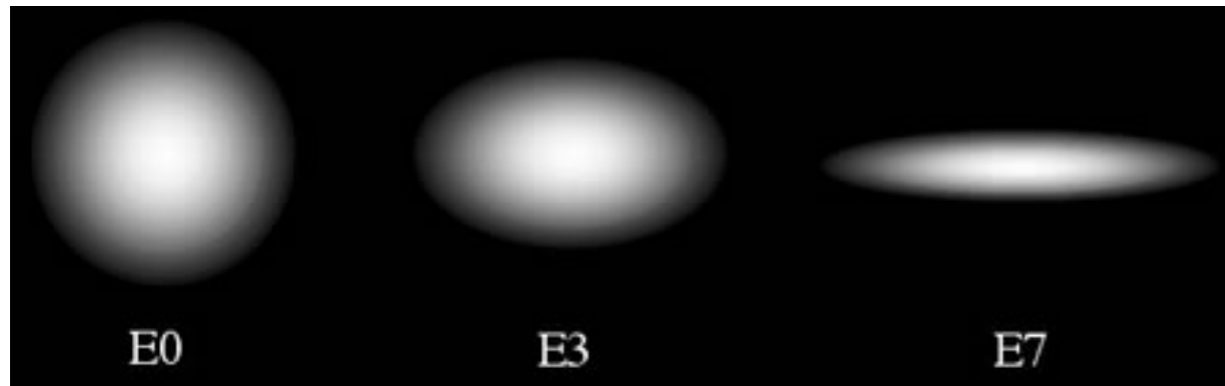
AGNs with following source density and emission power has right ingredients to maintain observed  $\text{CR}_{\text{EG}}$  density

$$1 \times 10^{-7} \text{ Mpc}^{-3}$$

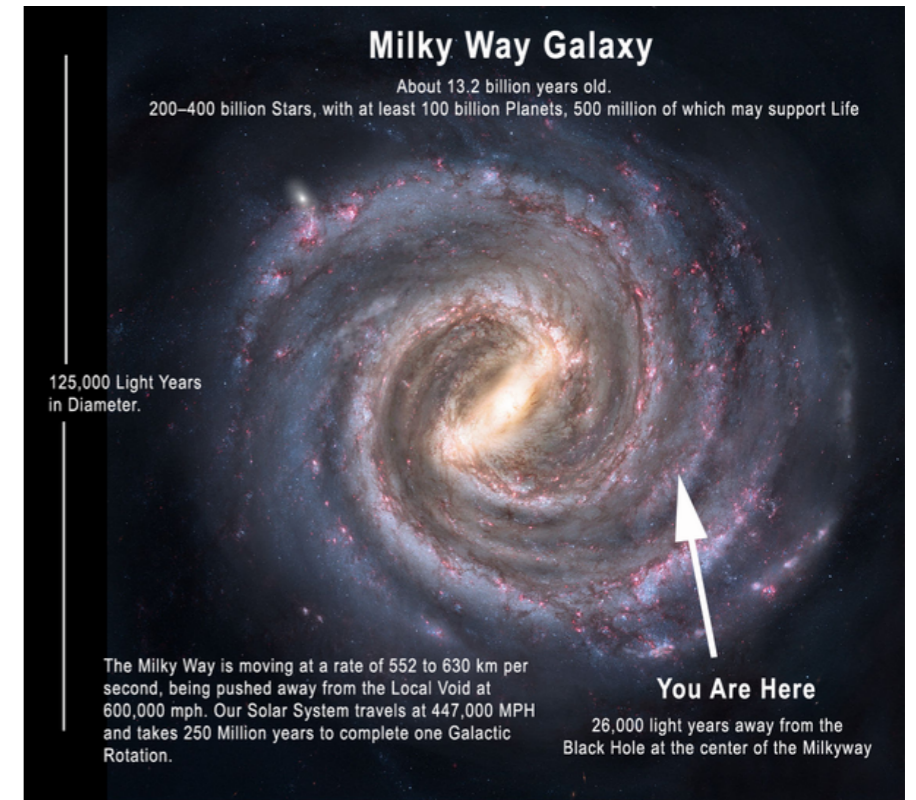
$$10^{44} \text{ erg/s}$$

# Normal Galaxy

Characterised by **Thermal Emission**



**Elliptical Galaxy**



**Spiral Galaxy**

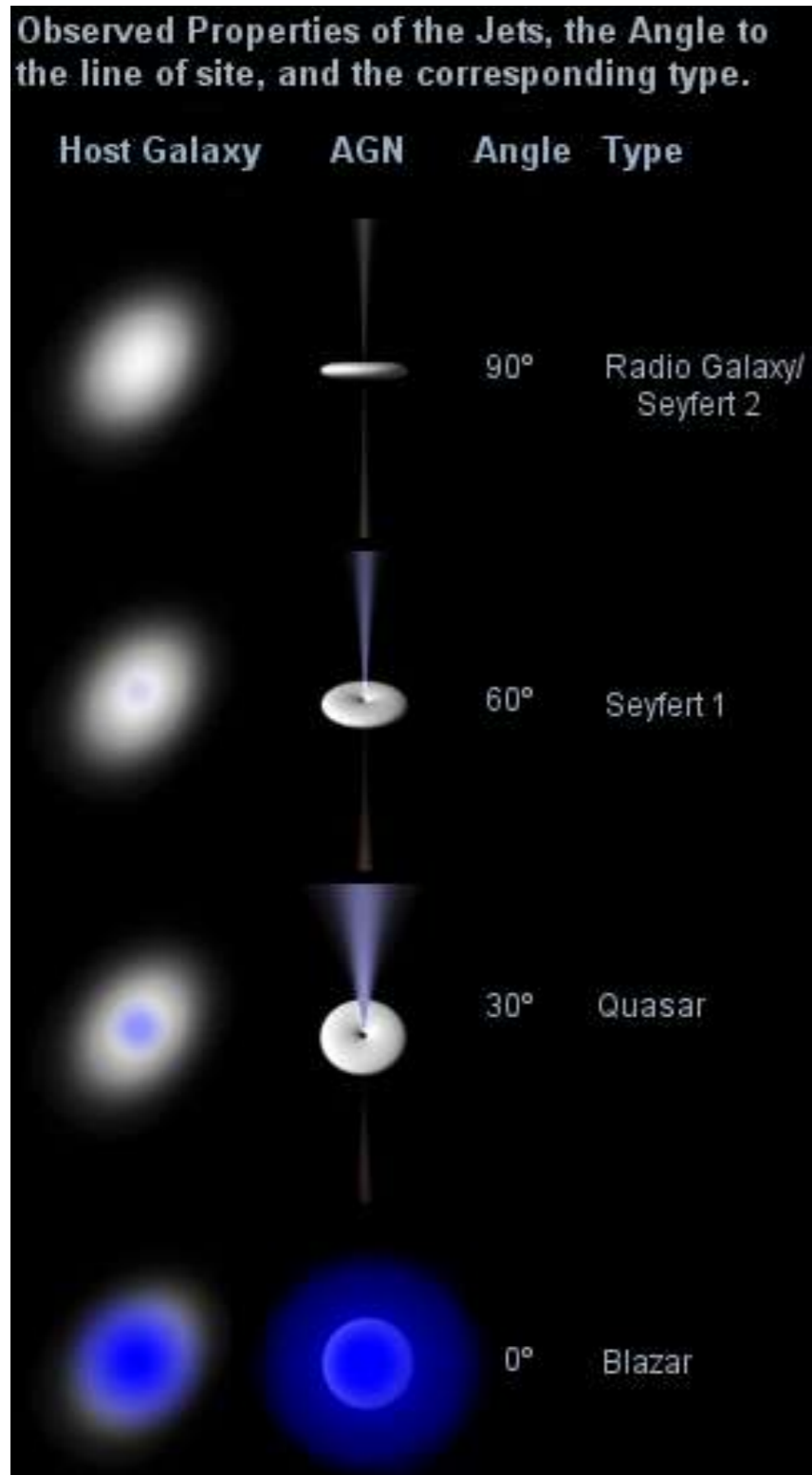
Normal Galaxies has total mass :  $10^7 - 10^{12}$  solar masses. Contains 100 billion stars.  
Luminosity  $10^8$  to  $10^{11}$  times the Sun.

Our Milky Way is a Spiral Galaxy. Diameter 50 kpc. Thickness 200 pc.  
 $1 \text{ pc} = 3.26 \text{ light-year or } 30 \text{ trillion kms}$

Solar System is located almost 27000 light-years from the galactic center.

# Active Galaxy

Characterised by **Non Thermal Emission**



- Very Bright , brightness  $\gg$  luminosity of normal galaxies
- Detected across EM spectrum from radio to gamma-rays
- Powered by accretion of mass onto a supermassive BH (Black Hole)  $10^9$ – $10^{11}$  solar masses

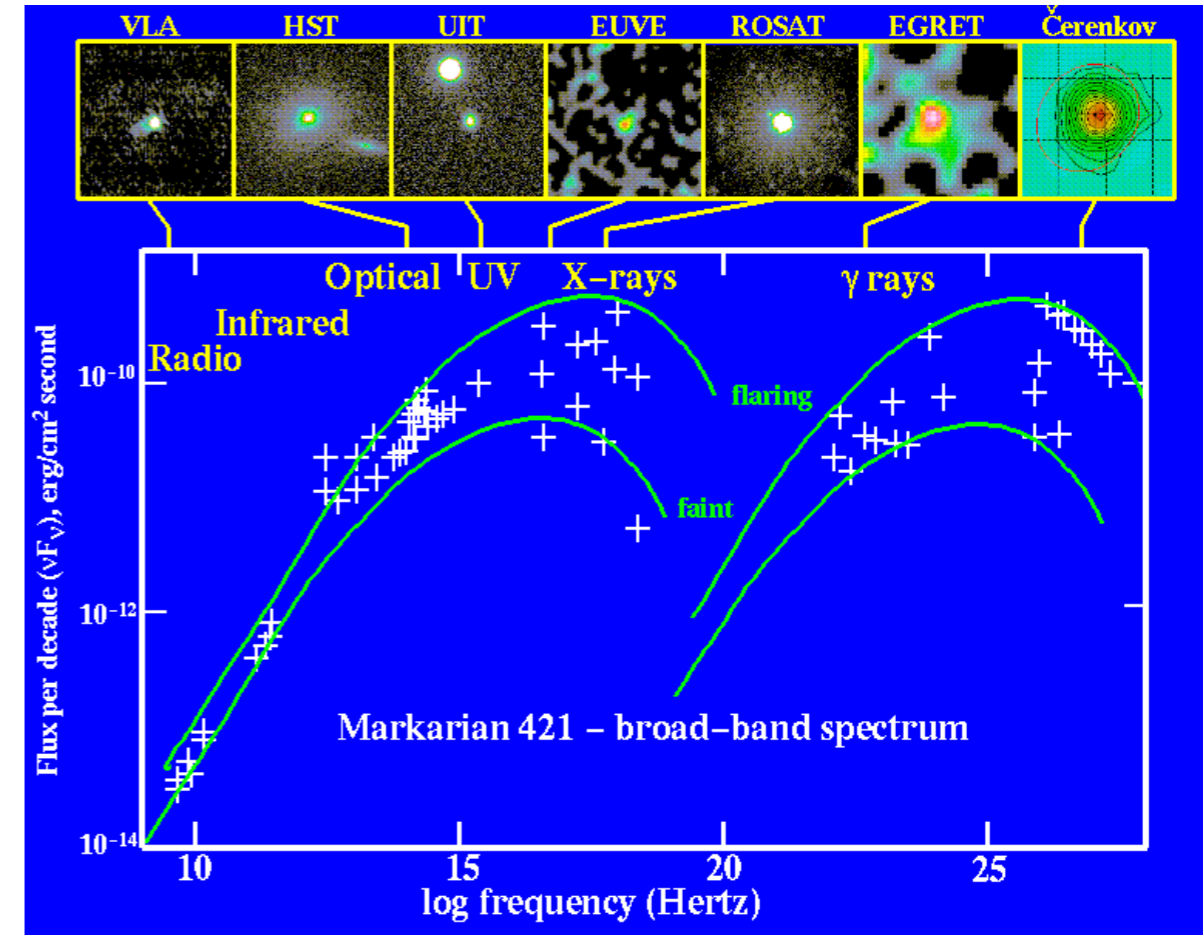
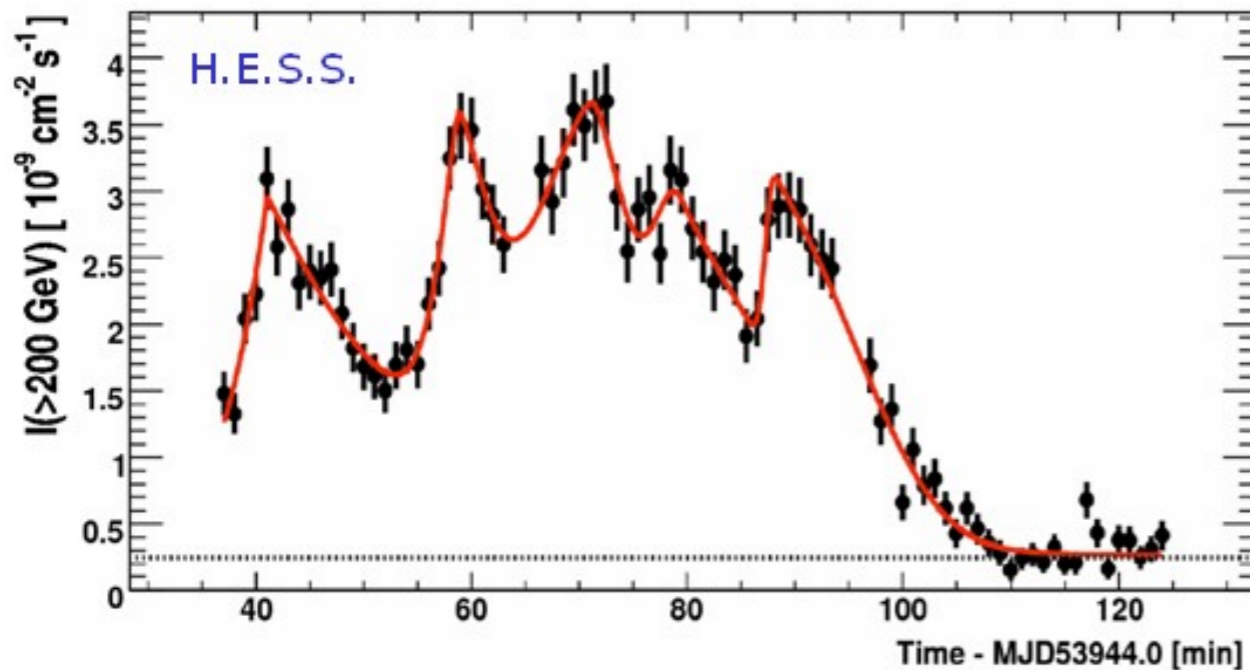
Radius SMBH :  $10^{13}$  cm

Radius of accretion disk :  $10^{14}$  cm

Extension of jet :  $10^{24}$  cm

**1-2% of all galaxies**

# Blazar a class of AGN



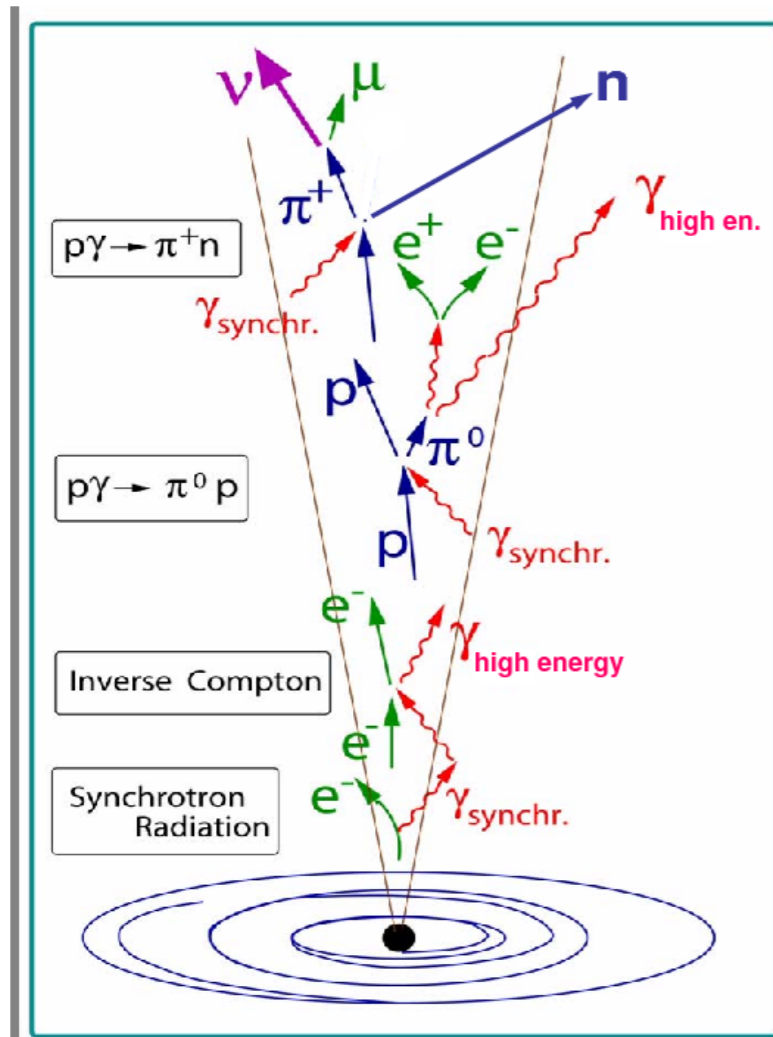
They are extremely interesting astrophysical laboratories – studying physical laws under relativistic conditions

Excellent probe of the Early Universe and everything between us and them

Less than 5% of all AGN

Origin of cosmic rays above  $10^{19}$  eV

# Emission Mechanisms @ Very High Energies



## Leptonic Models

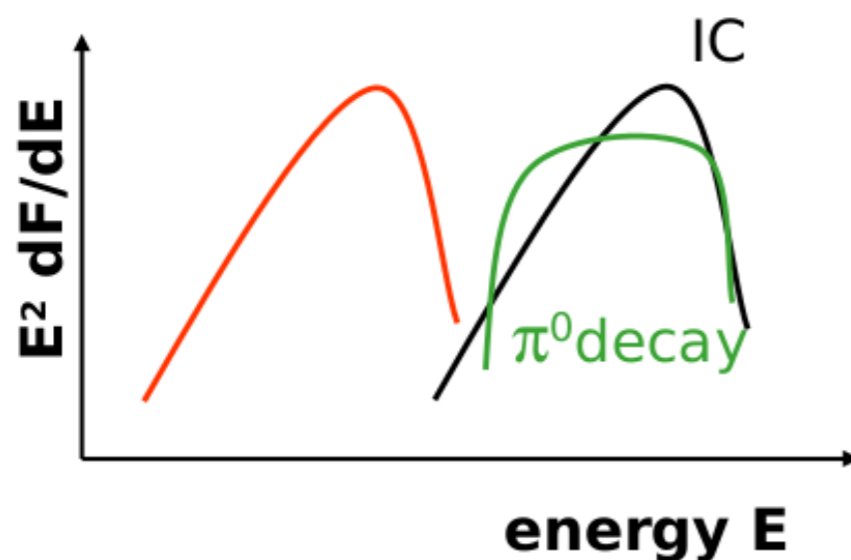
Synchrotron-Self Compton  
External Compton

## VHE $\gamma$ -rays from Blazars

Space-borne and ground based telescopes has detected VHE  $\gamma$ -rays :  
can be explained by leptonic models  
If they are source of cosmic rays  
then protons are also accelerated

## Hadronic Models

Proton-Proton Collisions  
Proton-Photon Collisions  
By Product **Neutrinos**

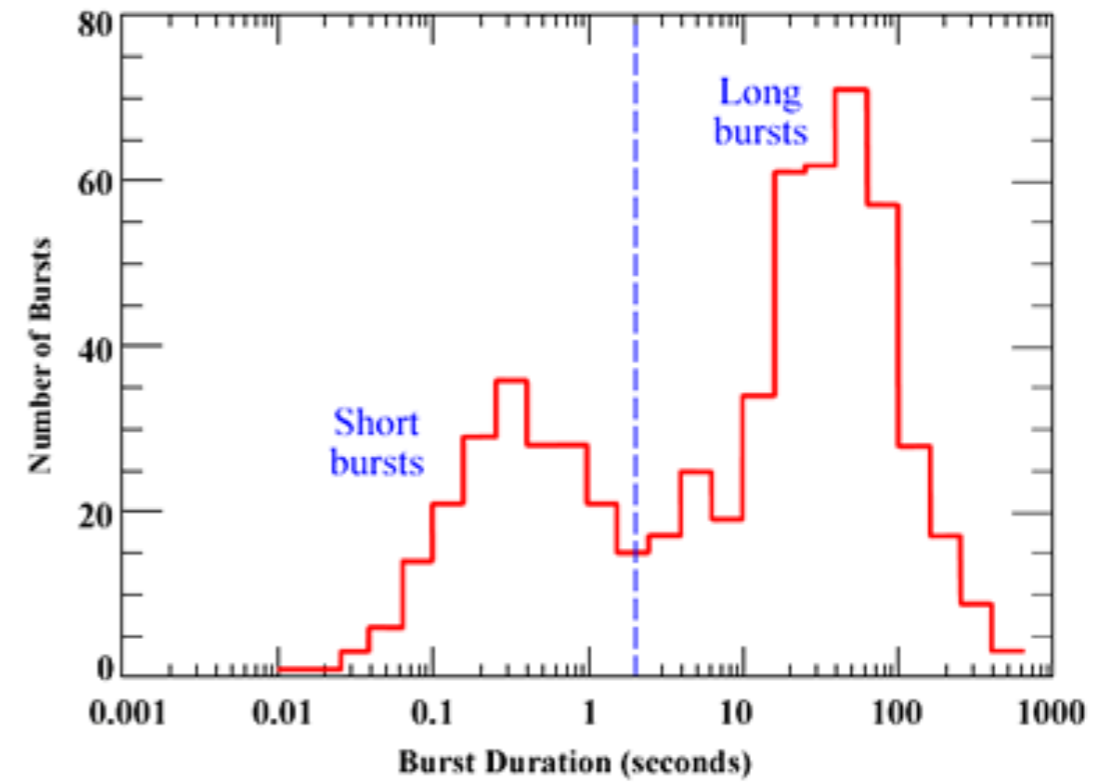
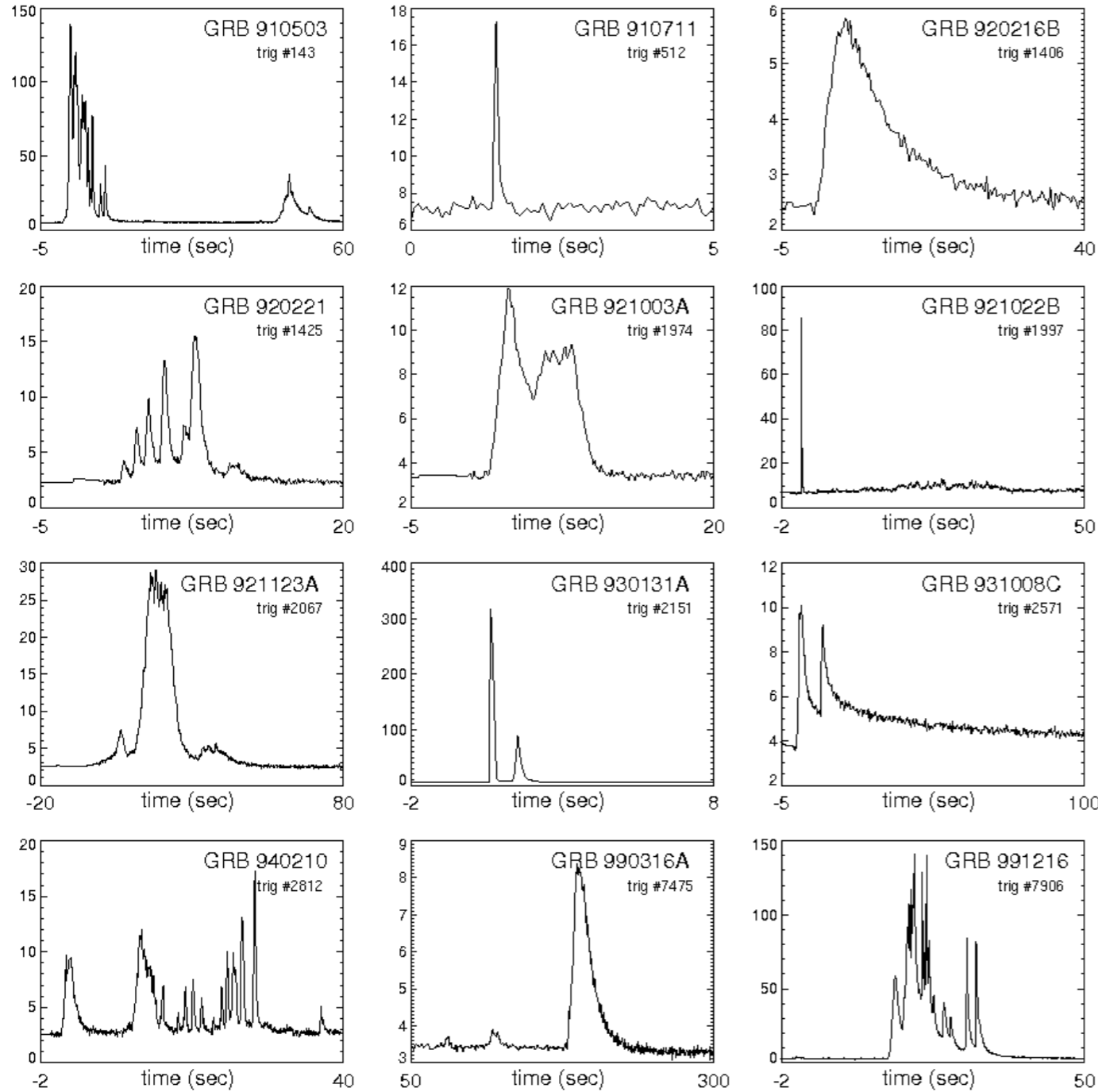


# Gamma Ray Burst (GRB)



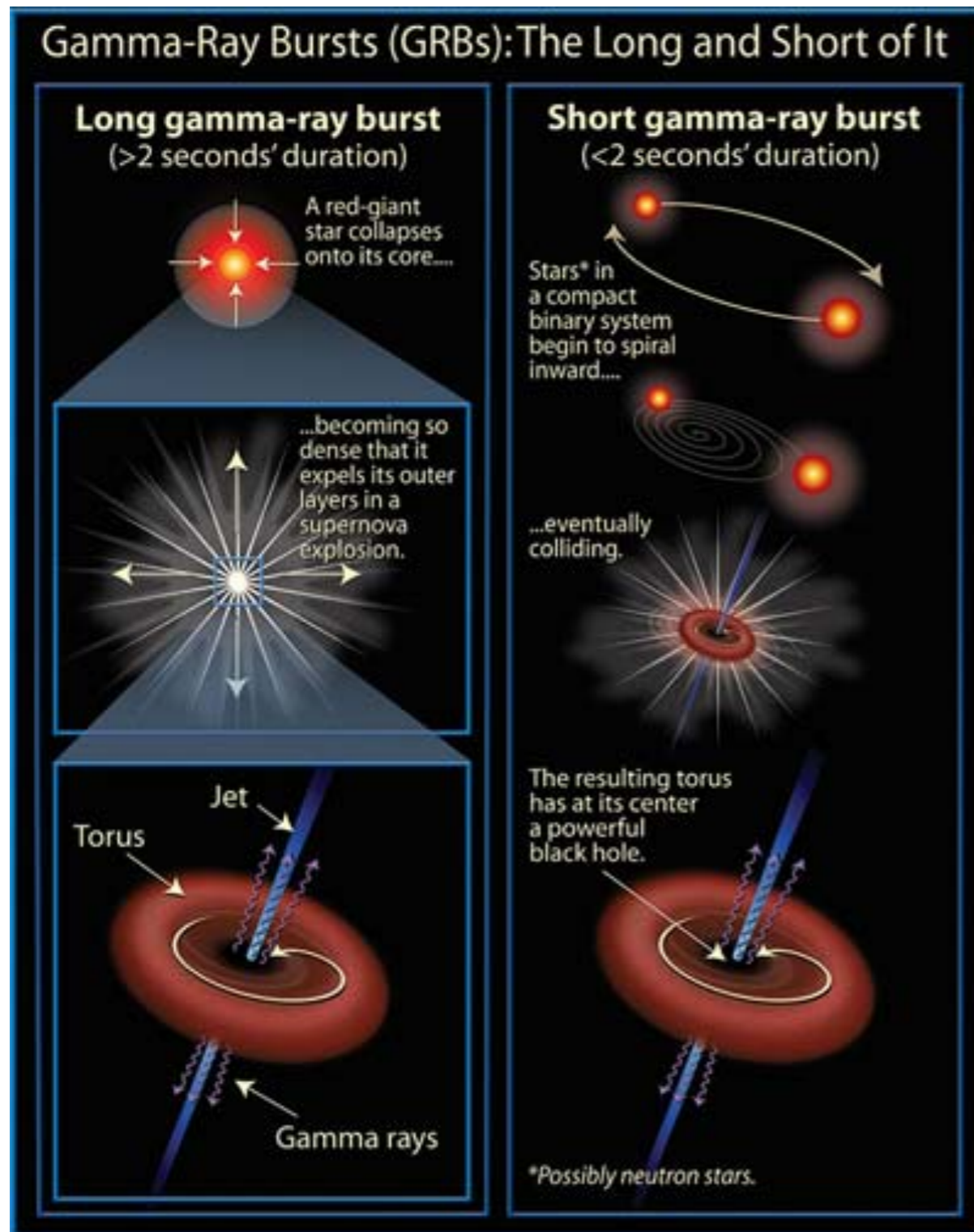
- GRBs are sudden, intense and non-repetitive flashes of gamma-rays
- They are the most luminous, most energetic events in the Universe, outshines everything else
- They lasts only for few seconds and release  $10^{54}$  ergs; comparable to SN, comparable to optical light emission by Sun over the lifetime of the Universe
- Non-thermal emission

# Variability of GRBs



**Rapid Variability and short timescales implies compact region and relativistic motion**

# Long & Short GRBs



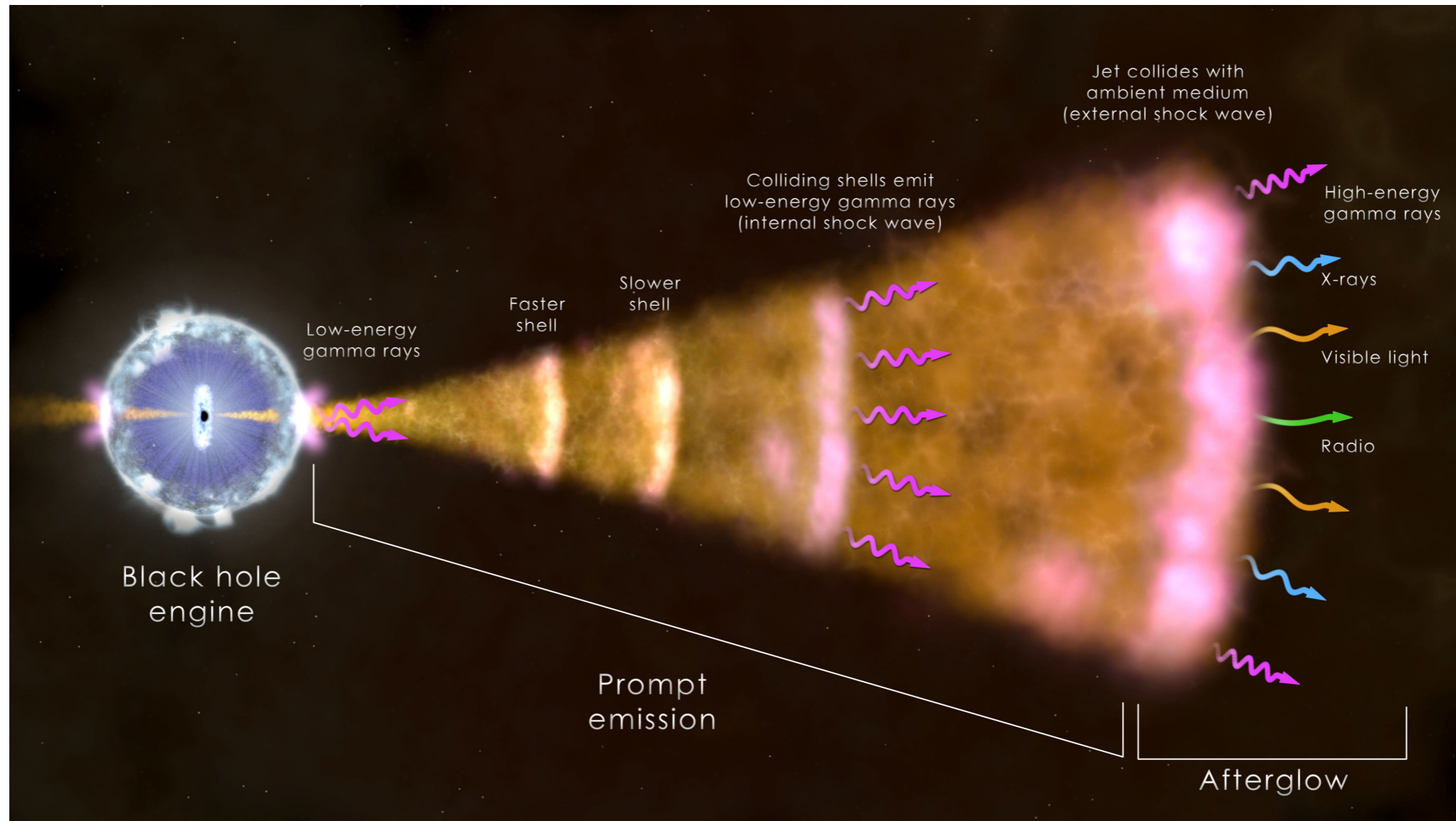
Upto 30 years after discovery only "prompt" emissions were detected from GRBs. Thus Location of these objects and nature of compact objects not known

Observation of afterglow changed that scenario

Long GRBs - core of massive star collapse [Evidence - absorption due to Fe in x-ray continuum spectrum.. associated with SN]

Short GRBs - Merger of NS-NS, NS-BH [Never associated with SN; found in elliptical galaxies- compact binaries most abundant]

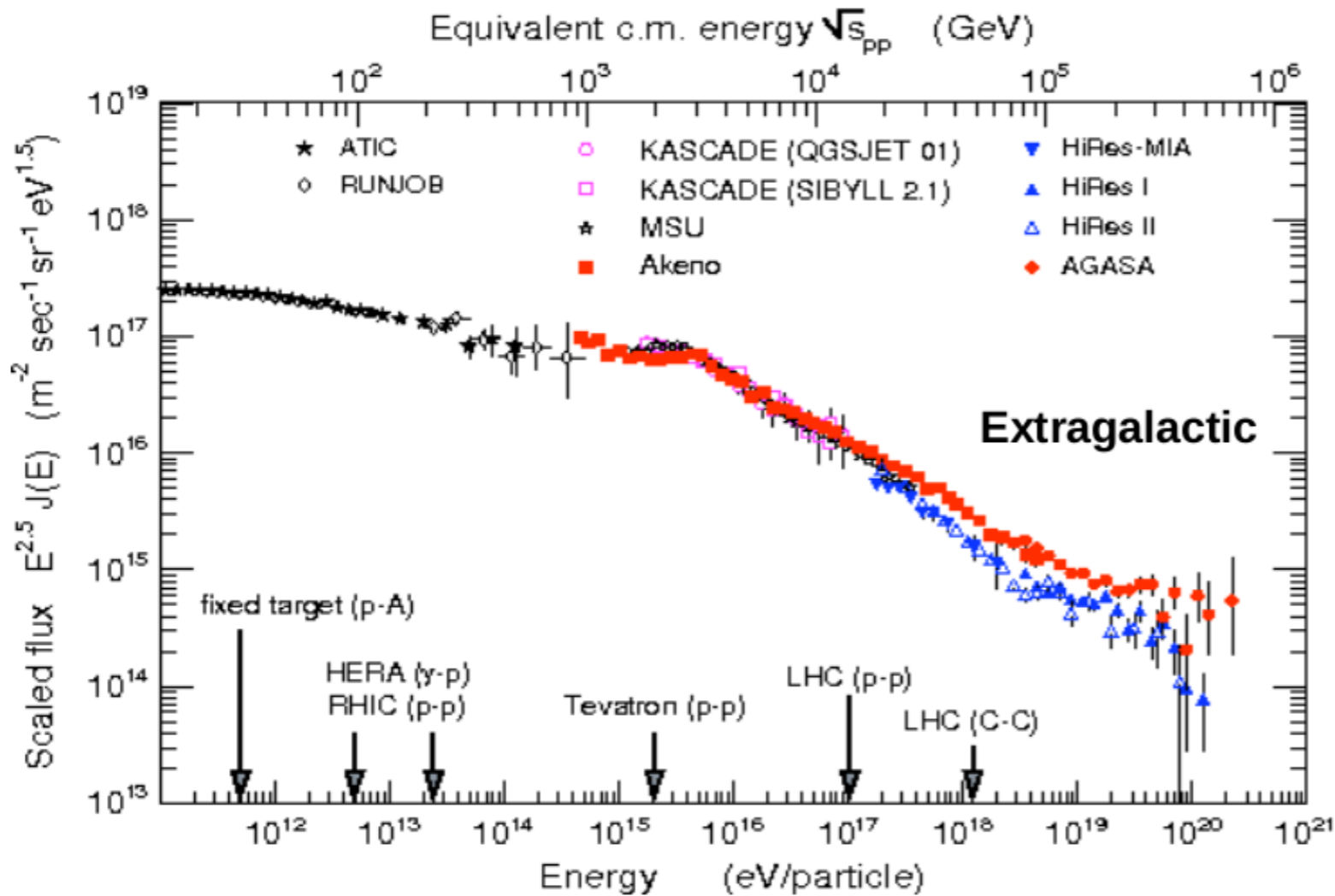
# Fireball Mechanism



The enormous energy release in a such a short time from very compact region produces tremendous luminosity which overwhelms gravity.

This outward pressure flings out matter, which gets heated up into a fireball made of electrons, positrons, gamma-rays, protons, neutrons

# Models Suggested GRBs only Source of VHE CRs



Energy Density of EG CR

$$\rho_{CR} \sim 10^{45} \text{ ergs/yr/Mpc}^3$$

Gamma-ray Bursts

$$\rho_{GRBs} \sim 10^{45} \text{ ergs/yr/Mpc}^3$$

From :

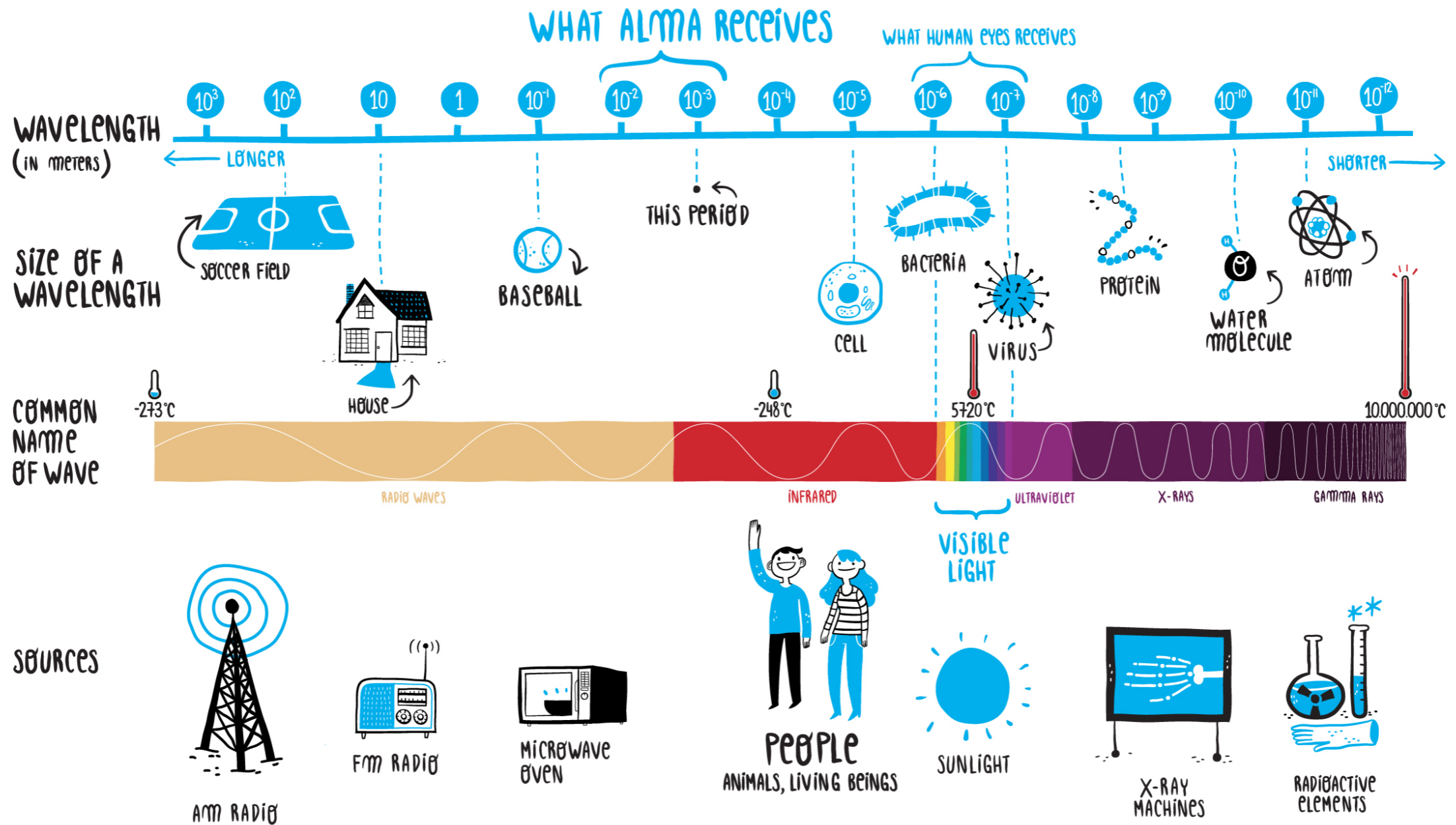
$$(2-3)10^{51} \text{ ergs} \times 300/\text{yr} / \text{Gpc}^3$$

**This Possibility is ruled out by IceCube Neutrino Telescope**

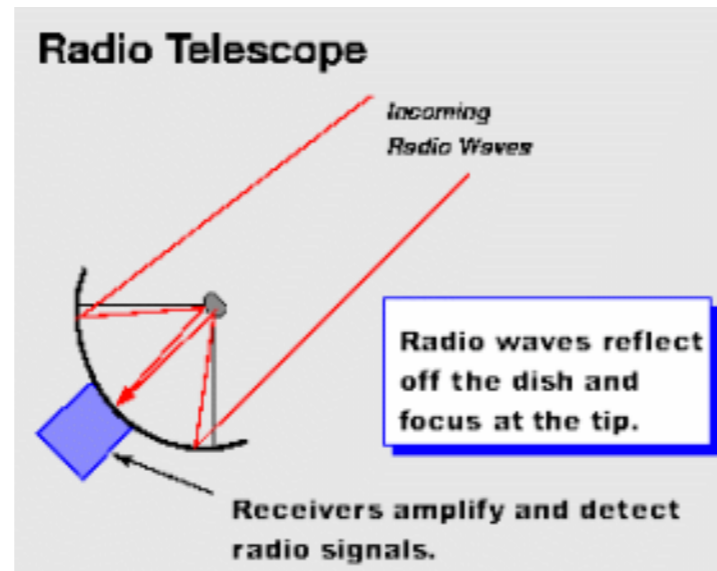
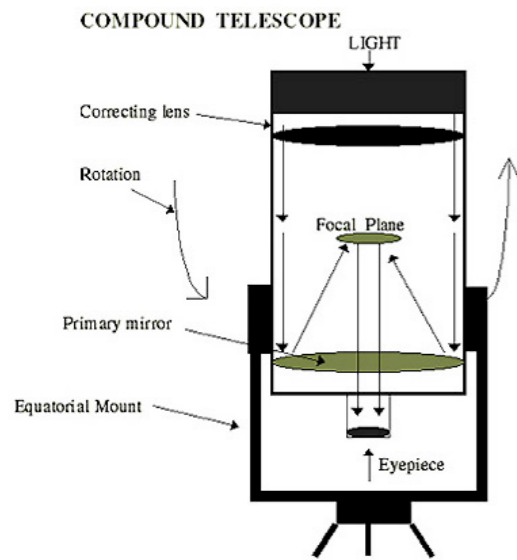
# How to Detect Gamma-rays & Neutrinos From Astrophysical Sources

# Gamma Ray Astronomy

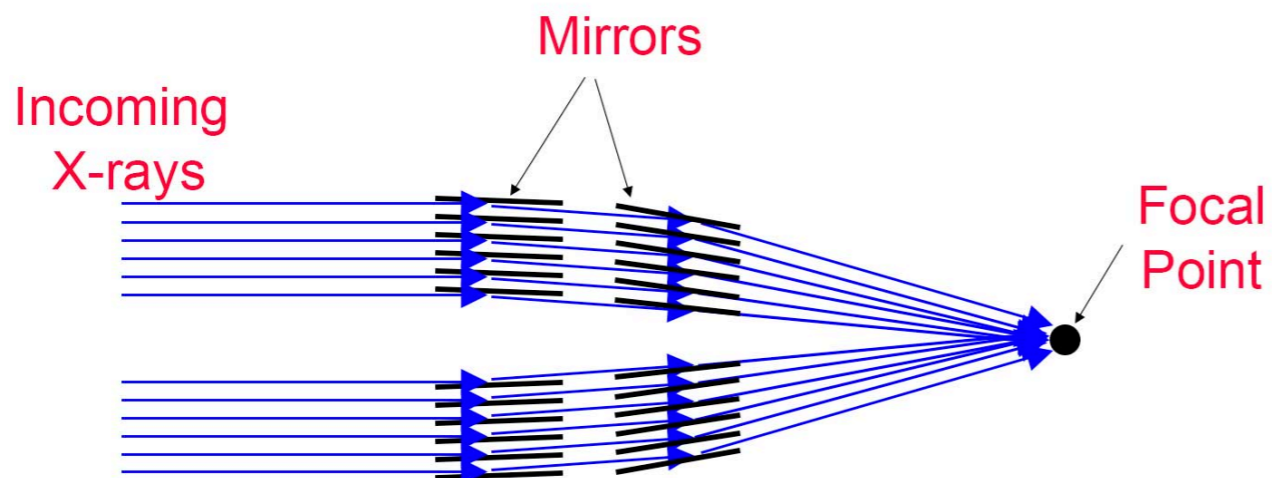
## THE ELECTROMAGNETIC SPECTRUM



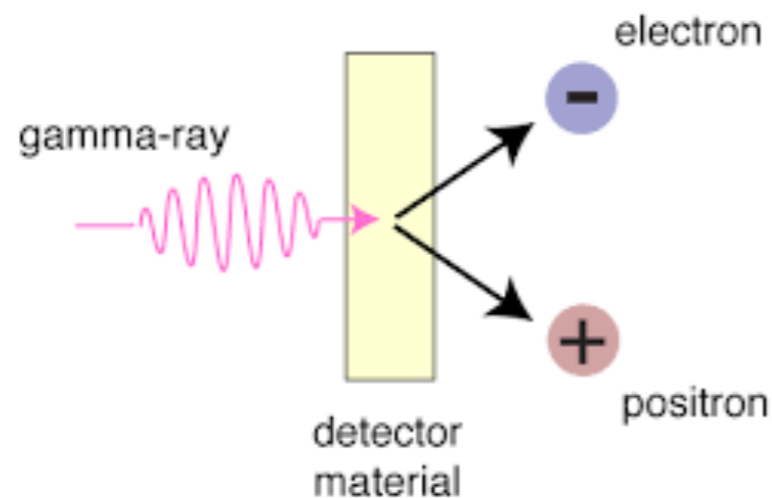
# Comparison : Optical/X-ray/Gamma-ray



Optical/radio wave is focused using a reflector : Reflection is caused by e-s reacting to EM field oscillating with same frequency and emits radiation with same frequency



X-rays can not be reflected like visible light, as index of refraction is close to 1 focal length becomes impractically long. However they can be reflected at shallow angle.

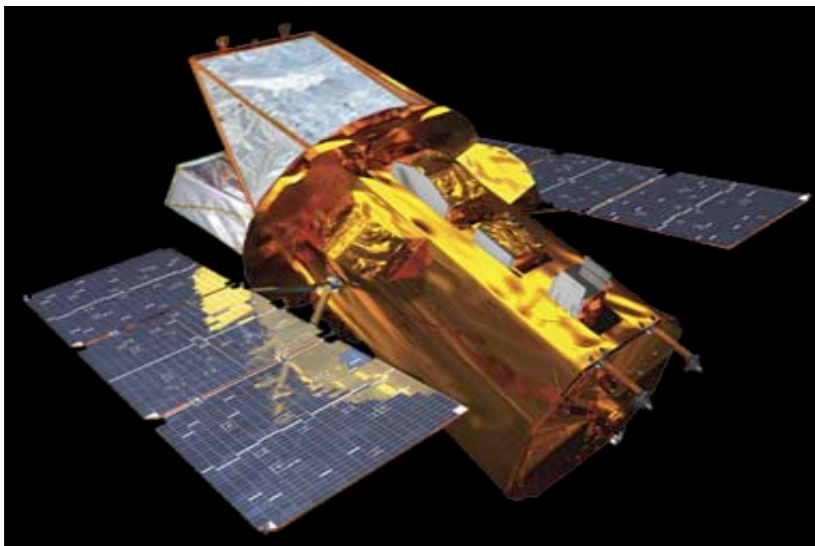


None of the above works for gamma-rays as they have very small wavelength they simple passes through matter. So they can be detected by pair production.

# Resolution : Optical/X-ray/Gamma-ray



Hubble Space telescope (Visible light) -  
1/20th of arc second

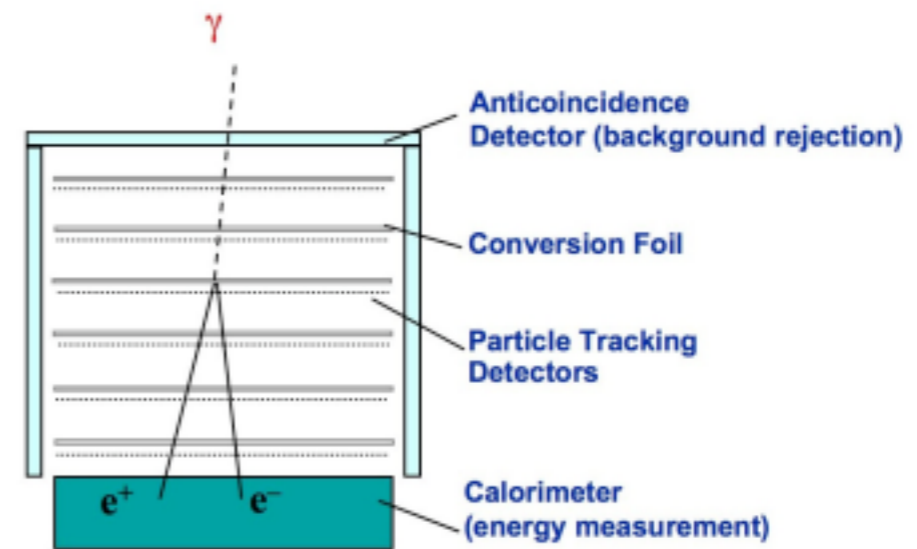
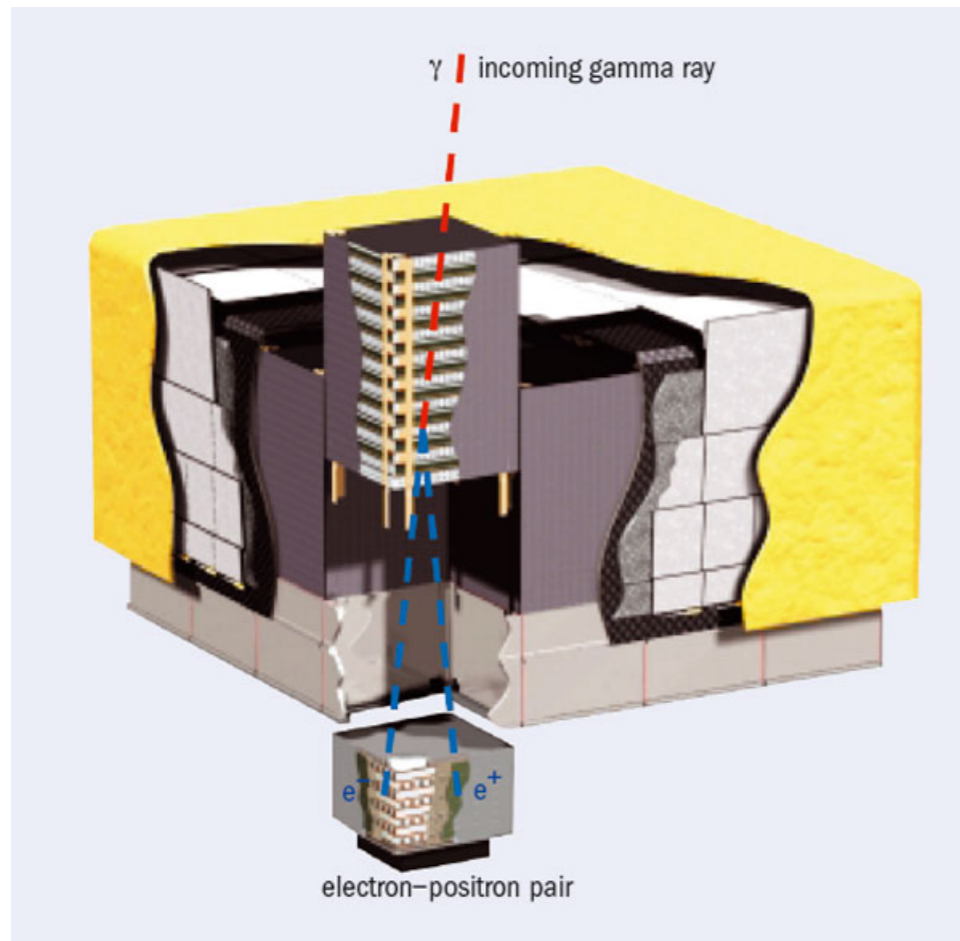
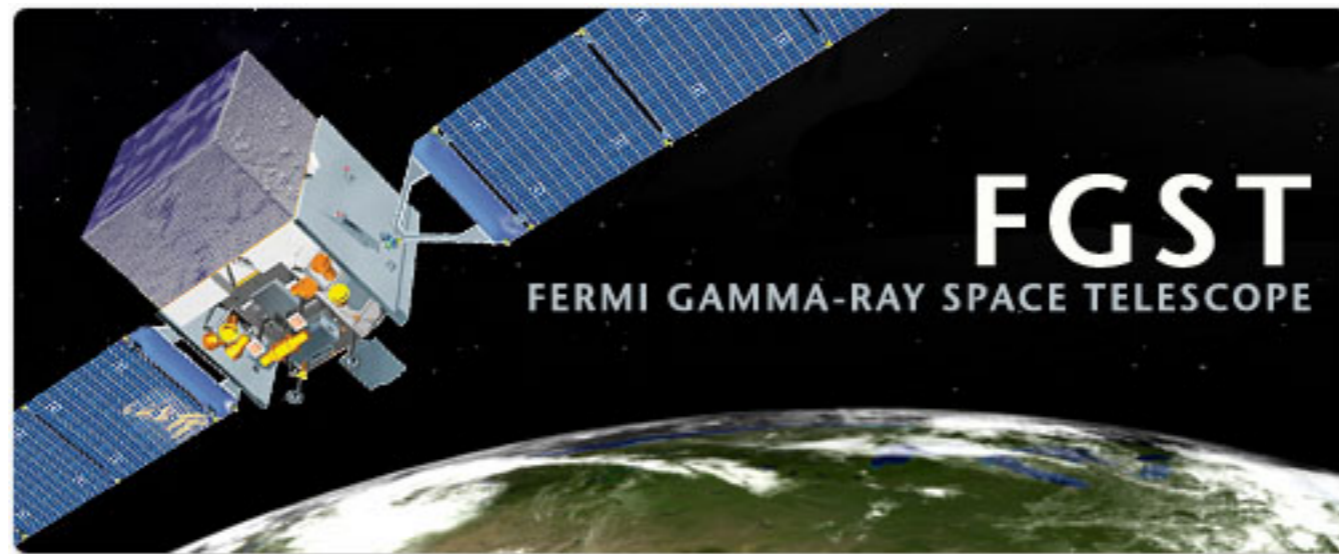


Swift X-ray telescope - 18 arc second



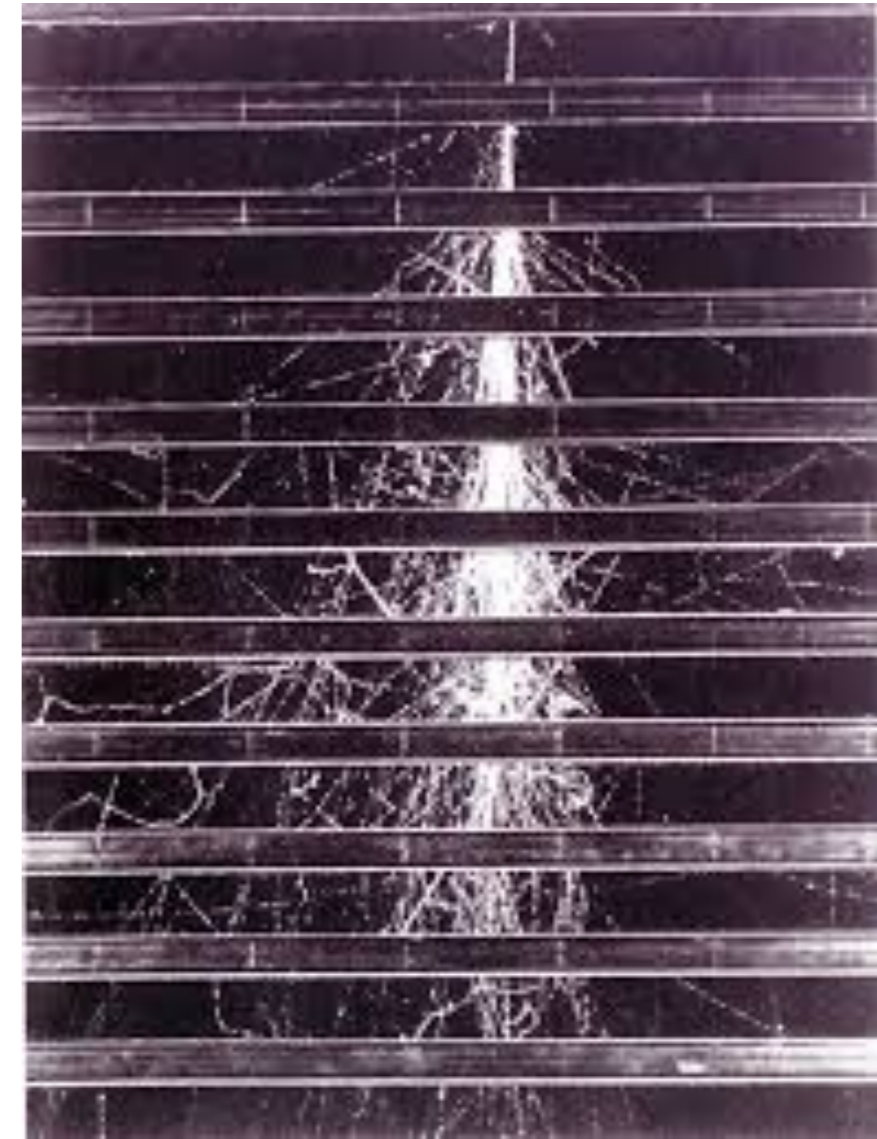
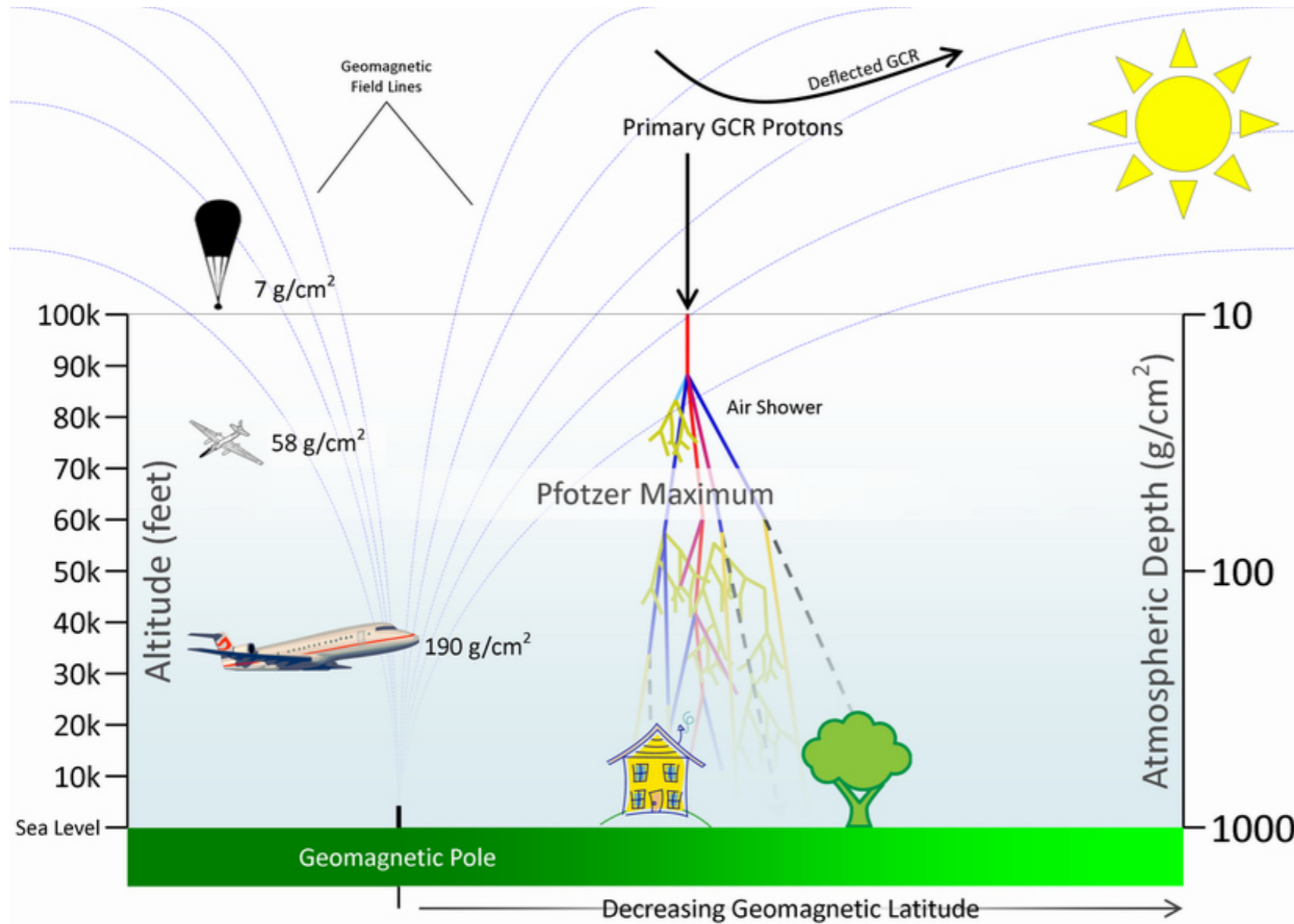
Fermi Gamma-ray telescope -  $1^\circ$

# Fermi Gamma ray Space Telescope



Area  $1\text{m}^2$   
Energy range MeV to GeV

# Gamma-ray Detection Technique @ Ground



**~130 cm of Iron**